The General Theory of Relativity

The general theory of relativity, Einstein's theory of gravitation, has been included as a compulsory subject in undergraduate and graduate courses in Physics and Applied Mathematics all over the world. However, the physics-first approach that is taken by many textbooks is not universally used, as the approach often depends on the instructors' or students' background. Conceived from the lecture notes made by the author over a teaching career spanning 18 years, this book introduces the general theory of relativity for advanced students with a strong mathematical background.

The proposed book takes a 'math-first approach', for which the mathematical formalism comes first and is then applied to physics. It presents a concise yet comprehensive and structured understanding of the general theory of relativity. The book discusses the mathematical foundation of the general theory of relativity and focuses heavily on topics such as tensor calculus, geodesics, Einstein field equations, linearized gravity, Lie derivatives and their applications, the causal structure of spacetime, rotating black holes, and basic knowledge of cosmology and astrophysics. All of these are explained through a large number of worked examples and exercises.

Farook Rahaman is a Professor of Mathematics at Jadavpur University, Kolkata. Besides writing a book, *The Special Theory of Relativity*, he has published numerous research papers on galactic dark matter, wormhole geometry, charged fluid model, topological defects in the early universe, gravastars, black hole physics, star modeling, and the cosmological model of the universe.

The General Theory of Relativity

A Mathematical Approach

Farook Rahaman



Cambridge University Press 978-1-108-83799-6 — The General Theory of Relativity Farook Rahaman Frontmatter <u>More Information</u>

CAMBRIDGE UNIVERSITY PRESS

University Printing House, Cambridge CB2 8BS, United Kingdom

One Liberty Plaza, 20th Floor, New York, NY 10006, USA

477 Williamstown Road, Port Melbourne, VIC 3207, Australia

314 to 321, 3rd Floor, Plot No.3, Splendor Forum, Jasola District Centre, New Delhi 110025, India

79 Anson Road, #06 04/06, Singapore 079906

Cambridge University Press is part of the University of Cambridge.

It furthers the University's mission by disseminating knowledge in the pursuit of education, learning and research at the highest international levels of excellence.

www.cambridge.org Information on this title: www.cambridge.org/9781108837996

© Farook Rahaman 2021

This publication is in copyright. Subject to statutory exception and to the provisions of relevant collective licensing agreements, no reproduction of any part may take place without the written permission of Cambridge University Press.

First published 2021

Printed in India

A catalogue record for this publication is available from the British Library

Library of Congress Cataloging-in-Publication Data Names: Rahaman, Farook, author. Title: The general theory of relativity : a mathematical approach / Farook Rahaman.

Description: Cambridge, United Kingdom ; New York, NY : Cambridge University Press, 2021.

Includes bibliographical references and index.

Identifiers: LCCN 2020037664 (print) | LCCN 2020037665 (ebook) |

ISBN 9781108837996 (hardback) | ISBN 9781108936903 (ebook)

Subjects: LCSH: Relativity (Physics)

Classification: LCC QC173.55 .R334 2021 (print) | LCC QC173.55 (ebook) | DDC 530.11–dc23 LC record available at https://lccn.loc.gov/2020037664 LC ebook record available at https://lccn.loc.gov/2020037665

ISBN 978-1-108-83799-6 Hardback

Cambridge University Press has no responsibility for the persistence or accuracy of URLs for external or third party internet websites referred to in this publication, and does not guarantee that any content on such websites is, or will remain, accurate or appropriate.

Cambridge University Press 978-1-108-83799-6 — The General Theory of Relativity Farook Rahaman Frontmatter <u>More Information</u>

То

my parents Majeda Rahaman and Late Obaidur Rahaman and my son and wife Md Rahil Miraj and Pakizah Yasmin

Contents

List of Figures		xiii
List of Tables		xvii
Preface		xix
Acknowledgments		xxi
Chapter	1 Tensor Calculus — A Brief Overview	1
1.1	Introduction	1
1.2	Transformation of Coordinates	1
1.3	Covariant and Contravariant Vector and Tensor	2
1.4	Operations on Tensors	6
1.5	Generalized Kronecker Delta	9
1.6	The Line Element	11
1.7	Levi-Civita Tensor or Alternating Tensor	18
1.8	Christoffel Symbols	20
1.9	Affine Connection	22
1.10	Covariant Derivative	24
1.11	Curvature Tensor	27
1.12	Ricci Tensor	29
1.13	Ricci Scalar	30
1.14	Space of Constant Curvature	32
1.15	The Affine Connection in Riemannian Geometry	36
1.16	Geodesic Coordinate	37
1.17	Bianchi Identity	38
1.18	Einstein Tensor	39
1.19	Weyl Tensor	41
Chapter	2 Geodesics	45
2.1	Geodesics Equation	45
2.2	Derivation of Euler–Lagrange Equation	46
2.3	Geodesic Equation in Curved Spacetime	47
2.4	Geodesic Deviation	49
2.5	Geodesics Are Auto Parallel	49
2.6	Raychaudhuri Equation	50

viii	C	ontent
Chapter	3 Einstein Field Equations	61
3.1	Introduction	6
3.2	Three Types of Mass	62
3.3	Einstein Tensor	62
3.4	Some Useful Variations	6.
3.5	Action Integral for the Gravitational Field	6.
3.6	Einstein's Equation from Variational Principle	64
3.7	Some Modified Theories of Gravity	70
Chapter	4 Linearized Gravity	8
4.1	Newtonian Gravity	8
4.2	Newtonian Limit of Einstein Field Equations or Weak Field Approximation	
	of Einstein Equations	8
4.3	Poisson Equation as an Approximation of Einstein Field Equations	90
4.4	Gravitational Wave	92
Chapter	5 Lie Derivatives and Killing's Equation	95
5.1	Introduction	9
5.2	Lie Derivative of a Scalar	90
5.3	Lie Derivative of Contravariant Vector	9
5.4	Lie Derivative of Covariant Vector	9′
5.5	Lie Derivative of Covariant and Contravariant Tensors of Order Two	9
5.6	Killing Equation	10
5.7	Stationary and Static Spacetimes	10
5.8	Spherically Symmetric Spacetime	10
5.9	Cylindrically Symmetric Spacetime (Axially Symmetry)	110
Chapter	6 Spacetimes of Spherically Symmetric Distribution of	
	Matter and Black Holes	11
6.1	Spherically Symmetric Line Element	11:
6.2	Schwarzschild Solution or Exterior Solution	11′
6.3	Vacuum Solution or Exterior Solution with Cosmological Constant	122
6.4	Birkhoff's Theorem	12.
6.5	Schwarzschild Interior Solution	12
6.6	The Tolman–Oppenheimer–Volkoff Equation	12
6.7	The Structure of Newtonian Star	129
6.8	Isotropic Coordinates	13
6.0		1.4

contents		ix
Chanter	7 Particle and Photon Orbits in the Schwarzschild Snacetime	159
7 1	Mation of Test Darticle	150
7.1	Function of Test Particle	139
7.2	Gravitational Redshift	101
7.3 7.4	Stable Circular Orbits in the Schwarzschild Spacetime	171
Chapter	8 Causal Structure of Spacetime	187
8.1	Introduction	187
8.2	Causality	187
8.3	Causal Relation	196
8.4	Causal Function	210
Chapter	9 Exact Solutions of Einstein Equations and Their Causal	
	Structures	219
9.1	Minkowski Spacetime	219
9.2	de Sitter Spacetime	225
9.3	Anti-de Sitter Space	230
9.4	Robertson–Walker Spaces	233
9.5	Penrose Diagrams of Robertson–Walker Spacetime for the Dust Case	235
9.6	Spatially Homogeneous Cosmological Models	237
9.7	Schwarzschild Solutions	240
9.8	Null Curves in Schwarzschild Spacetime	241
9.9	Time-like Geodesics in Schwarzschild Spacetime	242
9.10	Tortoise Coordinates	245
9.11	Eddington–Finkelstein Coordinates	245
9.12	Kruskal–Szekeres Coordinates	248
9.13	Reissner–Nordström Solution	253
Chapter	10 Rotating Black Holes	261
10.1	Null Tetrad	261
10.2	Null Tetrad of Some Black Holes	266
10.3	The Kerr Solution	271
10.4	The Kerr Solution from the Schwarzschild Solution	272
10.5	The Kerr–Newmann Solution from the Reissner–Nordström Solution	274
10.6	The Higher Dimensional Rotating Black Hole Solution	276
10.7	Different Forms of Kerr Solution	279
10.8	Some Elementary Properties of the Kerr Solution	283
10.9	Singularities and Horizons	284
10.10	Static Limit and Ergosphere	286
10.11	Zero Angular Momentum Observers in the Kerr Spacetime	288
10.12	Stationary Observer in the Kerr Spacetime	288
10.12	v 1	

292
293
295
297
301
305
305
307
309
312
313
315
316
318
319
320
322
324
326
328
334
335
335
337
338
339
341
342
345
345
348
350
357
360
361
364
366
366
370
371

Contents		xi
Appendix A	Extrinsic Curvature or Second Fundamental Form	379
Appendix B	Lagrangian Formulation of General Relativity	383
Appendix C	3+1 Decomposition	391
Bibliography Index		395 399

Figures

1	S and \overline{S} frames.	2
2	Two neighboring points in a space.	12
3	Locally Euclidean space.	22
4	Parallel transport.	23
5	Curves joining two fixed points.	46
6	Geodesic deviation.	49
7	The angle θ between A^i and t^i is constant.	59
8	Sphere of uniform mass.	87
9	Phenomenological comparison of Einstein and Newtonian theories.	90
10	Two neighboring points under infinitesimal one parameter transformation.	96
11	Direction of Killing vector along time axis.	109
12	Spherical symmetry.	110
13	Spherically symmetric body with uniform density.	134
14	Proper radial distance in Schwarzschild spacetime for $m = 1$ and $r_0 = 2$.	144
15	The embedding diagram for Schwarzschild spacetime for $m = 1$ in the left panel	
	and in the right panel we provide the entire imagining of the surface created by the	
	rotation of the embedded curve about the vertical z axis.	145
16	Plots for g_{tt} of Reissner–Nordström black hole and Schwarzschild black hole.	151
17	Planets are moving around the sun in an elliptic orbit.	162
18	Precession of the perihelion of the planet.	165
19	Deviation of the light ray passing near the sun.	166
20	Light signal passing through the gravitational field of the sun, from earth to the	
	planet and back after being reflected from the planet.	168
21	Two observers are sitting with clocks. Observer 1 is sending radiation to observer 2.	171
22	(Left) Effective potential of massless particle against $\frac{r}{m}$ for different values of $\frac{h^2}{m^2}$.	
	Curve 1 for $\frac{h^2}{m^2} = 8$, Curve 2 for $\frac{h^2}{m^2} = 10$, Curve 3 for $\frac{h^2}{m^2} = 12$, Curve 4 for	
	$\frac{h^2}{m^2}$ = 16. Note that V has only one extremum point, which is maximum. (Right)	
	Effective potential of massive particle against $\frac{r}{m}$ for different values of $\frac{h^2}{m^2}$. Curve	
	1 for $\frac{h^2}{m^2} = 8$, Curve 2 for $\frac{h^2}{m^2} = 10$, Curve 3 for $\frac{h^2}{m^2} = 12$, Curve 4 for $\frac{h^2}{m^2} = 16$.	
	Note that V has both maximum and minimum points.	174
23	Chronological future and past.	188
24	Causal future and past.	189
25	q lies in future of p on the causal curve γ .	190
26	Chronological (causal) future set of a set.	190

XIV		Figures
27	Every point in $I^+(n)$ is an interior point	190
28	All the limit points of $I^+(p)$ are not contained in $I^+(p)$	191
20	Every point in $I^+(p)$ is not an interior point	192
30	$I^+(n) \subset \overline{I^+(n)}$	192
31	r is an interior point	192
32	An event <i>n</i> lies just inside the boundary of the chronological future $\dot{J}^+(S)$	193
32	A future set of a set is the union of $I^+(n)$	193
34	A chronal set	104
35	$I^+(n) \subset I^+(S)$	104
36	F dge of a set	105
37	Convex normal neighborhood N of r	105
38	Every open neighborhood of x intersects infinity many $\begin{cases} 1 \\ 1 \end{cases}$	195
30	Every open neighborhood of x intersects infinity many χ_n . Identifying $t = 0$ and $t = 1$ hypersurfaces	190
39 40	The neighborhood V of n is contained in Q	197
40	The light cone of \overline{a} is strictly larger than that of a	190
41	The light cone of g_{ab} is survey larger than that of g_{ab} .	190
42	-v) is future-directed. When a future directed nonspace like curve leaves V the limiting value of fig	199
43	when a future-unected honspace-like curve leaves V, the mining value of J is greater than when a future directed nonspace like curve enters V	200
4.4	Eventure and past domain of domandance.	200
44	Couchy surface	201
43	Euture couchy horizon	202
40	Future cauchy horizon. No two points in $H^+(S)$ are time like related	202
47	No two points in $H^{+}(S)$ are time-like related.	203
40	$p \in D^{+}(S) - \Pi^{+}(S)$. The fature dense is a dense dense $D^{+}(S)$ and Cauchy having $U^{+}(S)$.	204
49 50	The future domain of dependence $D^{+}(S)$ and Cauchy horizon $H^{+}(S)$.	205
50	S is asymptotically null to the right and becomes exactly null to the feft.	203
51	The surface S_t of constant time in Minkowski spacetime.	206
52 52	S is not globally hyperbolic.	206
53	Example of nonglobally hyperbolic spacetime.	207
54	y lies strictly on the light cone.	207
33 56	One can join z and y through a time-like curve. The set $f = 1$	208
50	The point y lies outside of the causal future and past of $x \in M$.	208
57	Reflecting spacetime.	209
58	I'(q) is strictly contained in $I'(p)$.	211
59	$I^+(y) - I^+(x)$ contains an open set.	212
60	$I^+(y)$ is strictly contained in $I^+(x)$.	213
61	For future-directed time-like curve γ with a future end point p , $I^{-}(\gamma) = I^{-}(p)$.	215
62	The shaded region in figure is the TIP representing the point <i>p</i> .	215
63	The time-like geodesics γ from p .	217
64	Null coordinate $v(w)$ can be regarded as an incoming (outgoing) spherical wave.	221
65	(t, r) and (v, w) in a single origin.	221
66	Einstein static cylinder can be decomposed into various components.	223
67	Diagram of Minkowski spacetime in (t', r') plane.	224
68	Any point can be causally connected with future time-like infinity.	225

ures		>
69	The image of de Sitter spacetime. This is a hyperboloid embedded in a flat five-	
	dimensional spacetime given by general coordinates (t, χ, θ, ϕ) .	2
70	Einstein static universe.	2
71	Past infinity (I^-) and future infinity (I^+) , which are S^3 sphere.	2
72	Particle horizon.	2
73	In Minkowski spacetime all the particles are seen at any event p on $O's$ world.	2
74	An accelerating observer <i>R</i> in Minkowski space may have future event horizon.	2
75	Future and past event horizons.	2
76	Penrose diagram in anti-de Sitter space.	2
77	Robertson–Walker spacetime for $k = 1$ is mapped into the region in the Einstein static universe.	
78	Past infinity (I^{-}) and future infinity (I^{+}) in Robertson–Walker spacetime for $k = 1$.	2
79	Penrose diagram of Robertson–Walker spacetime for $k = -1$.	2
80	Dust-filled Bianchi-I model in $\tau - \eta$ plane.	ź
81	Outgoing and ingoing radial null geodesics.	ź
82	If we go toward $r = 2m$, the light cones become thinner and thinner and ultimately collapse entirely.	,
83	(Left) A body takes finite proper time to reach from $r = 2m$ to $r = 0$. (Right) Any	
	time-like particle requires infinite amount of time to touch the surface $r = 2m$.	2
84	The light cones in Schwarzschild geometry for the Tortoise coordinate (r^*, t) .	
85	The behavior of the light cone in Eddington–Finkelstein coordinate.	
86	The behavior of the light cone in Eddington–Finkelstein coordinate.	
87	Kruskal–Szekeres diagram.	
88	Penrose diagram of Schwarzschild solution in Kruskal coordinates.	
89	Penrose diagram of naked singularity in Reissner–Nordström solution.	
90	Light cones in Reissner–Nordström sacetime.	
91	Light cones in Reissner–Nordström spacetime.	
92	Light cones in extreme Reissner–Nordström spacetime.	
93	The figure indicates the position of the horizons, ergosurfaces, and curvature	
	singularity in the Kerr black hole spacetime.	ć
94	Light cones in Kerr spacetime.	ć
95	Conformal structure of the Kerr spacetime.	,
96	Conformal structure of the extreme Kerr spacetime.	ć
97	Hawking radiation.	2
98	Penrose process.	4
99	Particles are moving along nonintersecting geodesics.	
100	The sides of a triangle are expanded by the same factor.	2
101	Galaxy has linear extend $d(\overline{AB})$.	
102	Universe with $k = 0, -1, 1$.	-
103	Behavior of $a(t)$ in de Sitter model.	-
104	Photoelectric method.	-
105	H-R diagram.	-
106	Chemical compositions of stars	2

xvi		Figures
107		240
107	Internal structure of the star.	348
108	Fusion creates an external pressure that stabilizes the inward pressure caused by	T
	gravity, steadying the star.	360
109	Collapsing star.	369
110	Gravitational lensing diagram.	369
111	3+1 foliation of spacetime: Decomposition of t^{α} into lapse and shift.	392

Tables

Ι	Theoretical prediction and observed values of the advance of perihelion of some	
	planets	165
II	Solution of Lane–Emden equation.	354

Preface

At the beginning of the twentieth century, Einstein spent many years developing a new theory in physics. The newly developed theory is known as the theory of relativity. This is basically a combination of two theories: the first one is known as the special theory of relativity and latter one is dubbed as the general theory of relativity. The special theory of relativity is based on two postulates, namely the principle of relativity or equivalence, that is, the laws of physics are the same in all inertial systems, which means no preferred inertial system exists, while the second postulate is the principle of the constancy of the speed of light. The general theory of relativity asserts that there is no difference between the local effects of a gravitational field and that of acceleration of an inertial system. In other words, spacetime is warped or distorted by the matter and energy in it as an effect of gravity. According to the general theory of relativity, massive objects cause the outer space to twist due to gravity like a heavy ball bending a thin rubber sheet that is holding the ball. Heavier balls bend spacetime far more than lighter ones. Like the special theory of relativity, the general theory of relativity attracted scientists a lot, immediately after its discovery by Einstein. As a result, it has been included as a compulsory subject in graduate and postgraduate courses of physics and applied mathematics all over the globe. Einstein proposed the field equations for the general theory of relativity by applying his own intuition. Later, many other methods were developed to construct Einstein's field equations.

This book on the general theory of relativity is an outcome of a series of lectures delivered by me, over several years, to postgraduate students of mathematics at Jadavpur University. I should mention that it is not a fundamental book. This book has been written, from a mathematical point of view, after consulting several books existing in the literature. I have provided the list of the reference books. During my lectures, many students asked questions that helped me know their needs as well as the shortcomings in their understanding. Therefore, it is a well-planned textbook that has been organized in a logical order and every topic has been dealt with in a simple and lucid manner. A number of problems with hints, taken from the question papers of different universities, are included in each chapter.

The book is organized as follows:

In Chapter One a brief overview of tensor calculus, including the different types of tensors as well as operations on tensors, is given. Generalized Kronecker delta, Christoffel symbols, affine connection, covariant derivatives, geodesic coordinate, and various forms of tensors are described, with examples, as a foreground to understand the basics of general relativity. Chapter Two starts with a discussion of the geodesic equation in curved spacetime. In addition, several problems for different spacetimes are provided on geodesics. Chapter Three begins with the statement of three basic principles, namely Mach's principle, equivalence principle, and the principle of covariance. Next, the Einstein gravitational field equations are derived from the variational principle.

Cambridge University Press 978-1-108-83799-6 — The General Theory of Relativity Farook Rahaman Frontmatter <u>More Information</u>

хх

Also, in this chapter, the outline of some modified theories of gravity, such as f(R) theory of gravity, Gauss-Bonnet gravity, f(G) theory of gravity or modified Gauss-Bonnet gravity, f(T) theory of gravity, f(R,T) theory of gravity, Brans–Dicke theory of gravity, and Weyl gravity, are provided. A discussion on linearized gravity is given in Chapter Four. Newtonian limit of Einstein field equations or weak field approximation of Einstein field equations is derived. It is shown that Poisson's equation can be viewed as an approximation of Einstein field equations. A short mathematical description of gravitational wave is also provided. Chapter Five is dedicated to a short discussion on Lie derivatives and their applications. Killing equations and Killing vectors are also discussed with several examples. A short note on conformal Killing vector is also provided. Chapter Six is devoted to discussions on spacetimes of spherically symmetric distributions of matter. The exact exterior and interior solutions of Einstein field equations in spherically symmetric spacetimes are discussed. The proof of Birkoff's theory is provided. It states that a spherically symmetric gravitational field in vacuum is necessarily static and must have Schwarzschild form. The Tolman-Oppenheimer-Volkov (TOV) equation is discussed. Isotropic coordinate system is a new coordinate system whose spatial distance is proportional to the Euclidean square of the distances. Some static spherically symmetric spacetimes are rewritten in an isotropic coordinate system. A short discussion on interaction between the gravitational and electromagnetic fields are provided. Reissner–Nordström solution is a static solution of the gravitational field outside of a spherically symmetric charged body. Particle and photon orbits in the Schwarzschild spacetime are discussed in Chapter Seven. Also, in this chapter, using the trajectory in the gravitational field of sun (i.e., in the Schwarzschild spacetime), several tests of the theory of general relativity, namely the precession of the perihelion motion of mercury, bending of light, radar echo delay, and gravitational redshift, are explained. A discussion on the stable circular orbits in the Schwarzschild spacetime is given. A general treatment is provided for the experimental test of general theory of relativity for a general static and spherically symmetric configuration. Causal structure in the special theory of relativity, i.e., in Minkowski spacetime or flat spacetime, is characterized so that no massive particle can travel faster than light. In general relativity, locally there is no difference of the causality relation with Minkowski spacetime. However, globally, the causality relation is significantly different due to various spacetime topologies. A short discussion on causal structure of spacetimes is given in Chapter Eight. Several basic definitions and some standard theorems related to causality are explained. Chapter Nine deals with discussions on causal structures of specific spacetimes, which are the standard exact solutions of Einstein field equations such as Minkowski spacetime, de Sitter and anti-de Sitter spacetimes, Robertson-Walker spacetime, Bianchi-I spacetime, Schwarzschild spacetime, and Reissner-Nordström black hole. A short elementary discussion on rotating black holes is given in Chapter Ten. After introducing the tetrad, an outline of the derivation of the Kerr and Kerr-Newman solutions is illustrated through the complex transformation algorithm for both in four and higher dimensions. Some of the different forms of the Kerr solution are mentioned. Some elementary properties of the Kerr solution including the maximal extension of Kerr spacetime are discussed. Finally, brief discussions on Hawking radiation, Penrose process of extraction of energy from a Kerr black hole, and laws of black hole thermodynamics are given. Chapters Eleven and Twelve provide some simple applications of general theory of relativity in astrophysics and cosmology, respectively. Some preliminary concepts of extrinsic curvature, Lagrangian formalism of the general theory of relativity, and 3 + 1 decomposition of spacetime are given as appendices.

Acknowledgments

This book has been made possible through the support, contributions, and assistance of many people and various organizations. I take this opportunity to express my sincere gratitude to all of them. I would like to deeply and sincerely thank my mother (Majeda), wife (Pakizah), and son (Rahil), without whose loving support and encouragement this book could not have been completed. I also express my sincere gratitude to my father-in-law (Abdul Hannan), mother-in-law (Begum Nurjahan), brother-in-law (Dr. Ruhul Amin), younger brother (Mafrook Rahaman), and niece (Ayat Nazifa) for their patience and support during the entire period of the preparation of the manuscript. It is a pleasure to thank Dr. Nupur Paul, Dr. Sayeedul Islam, Dr. Banashree Sen, Dr. Mosiur Rahaman, Dr. Indrani Karar, Monsur Rahaman, Dr. Shyam Das, Lipi Baskey, Nayan Sarkar, Md Rahil Miraj, Dr. Arkopriya Mallick, Sabiruddin Molla, Dr. Ayan Banerjee, Dr. Tuhina Manna, Dr. Amna Ali, Dr. Nasarul Islam, Ksh. Newton Singh, Somi Aktar, Bidisha Samanta, Dr. Sourav Roychowdhury, Dr. Debabrata Deb, Dr. Amit Das, Dr. Abdul Aziz, Dr. Anil Kumar Yadav, Monimala Mandal, Antara Mapdar, Dr. Saibal Ray, Dr. Mehedi Kalam, Susmita Sarkar, Dr. Piyali Bhar, Dr. Gopal Chandra Shit, Dr. Ranjan Sharma, Dr. Shounak Ghosh, and Dr. Iftikar Hossain Sardar for their technical assistance in the preparation of the book. I remain thankful to all the professors and non-teaching staff members of the Department of Mathematics, Jadavpur University for providing me with all the available facilities and services whenever needed. Particularly I would like to mention the library staff for their excellent support. Finally, I am also thankful to the authority of the Inter-University Centre for Astronomy and Astrophysics (IUCAA), Pune, India for providing all kinds of working facility and hospitality under the Associateship Scheme.