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Introduction

As already mentioned above a main tool to identify and investigate *solvable* dynamical systems—including N -body models describing the evolution in the *complex* plane (or equivalently in the *real* Cartesian plane) of N point particles interacting *nonlinearly* among themselves and moving according to *Newtonian* equations of motion (“accelerations equal forces”)—are the relations among the *time-evolutions* of the N *zeros* and those of the N *coefficients* of a monic *time-dependent* polynomial of (*arbitrary*) degree N . This approach was introduced decades ago [16] and has been extensively used since—see, for instance, the two books [20] and [27] and references therein—but the technique to implement it was essentially restricted to the consideration of the time evolution of the *zeros* of a (monic) polynomial the *coefficients* of which evolve according to a system of *linear* Ordinary Differential Equations (ODEs). The elimination of this restriction—to *linear* evolutions of the *coefficients*—was recently made possible by the key formulas reported in Chapter 2, opening thereby the way to the identification of many more *solvable/integrable* N -body problems, as described in Chapter 4 and demonstrated by the examples reported there. This approach has also been extended to the identification and investigation of *solvable nonlinear Partial Differential Equations* (PDEs: see Chapter 5) and to equations of evolution in *discrete time* (see Chapter 7).

Moreover—as described in Chapter 6—this development naturally led to the idea of *generations* of monic polynomials of degree N —with the N *coefficients* of the polynomials of the *next* generation identified with the N *zeros* of the polynomials of the *current* generation—and to the related possibility to identify *endless hierarchies* of *nonlinear solvable* dynamical systems, including, again, N -body problems describing the evolution in the *complex* plane (or, equivalently, in the *real* Cartesian plane) of N point-particles interacting *nonlinearly* among themselves and moving according to *Newtonian* equations of motion (“accelerations equal forces”).

Some other developments that have also followed from the key formulas reported in Chapter 2 are treated in other chapters and sections of this book: the interested

reader may immediately identify them from the titles of the relevant chapters, sections and subsections, as displayed in the Table of Contents.

Let us end this terse Chapter 1 by displaying, without any immediate explanation, 2 *rather trivial examples* of the type of N -body problems—for simplicity, with $N = 2$ —the *solvability* of which, demonstrated in this book, entails an understanding of their *remarkable* phenomenology. Hence, in the rest of this short Chapter 1 we refer to 2 point particles moving in the *real* Cartesian plane, while in the rest of this book we mainly focus on N -body problems—with $N \geq 2$ —characterizing the motion of N points moving in the *complex* plane.

Example 1.1. The equations of motion of the *first* of these 2 *solvable* models read as follows:

$$\begin{aligned} \ddot{\vec{r}}_n = & (-1)^{n+1} \frac{2}{r^2} \left[\left(\dot{\vec{r}}_1 \cdot \dot{\vec{r}} \right) \dot{\vec{r}}_2 + \left(\dot{\vec{r}}_2 \cdot \dot{\vec{r}} \right) \dot{\vec{r}}_1 - \left(\dot{\vec{r}}_1 \cdot \dot{\vec{r}} \right) \dot{\vec{r}} \right] \\ & - (\rho_1 \omega)^2 \vec{r}_n + \left[2 (\rho_1)^2 - (\rho_2)^2 \right] \omega^2 r^{-2} \left[(r_{n+1})^2 \vec{r}_n - (r_n)^2 \vec{r}_{n+1} \right], \\ \vec{r} \equiv & \vec{r}_1 - \vec{r}_2, \quad n = 1, 2 \bmod [2]. \end{aligned} \quad (1.1)$$

Here the 2-vectors $\vec{r}_1 \equiv \vec{r}_1(t)$ and $\vec{r}_2 \equiv \vec{r}_2(t)$ are the coordinates of 2 unit-mass point-particles moving in the *real* Cartesian plane as functions of the *independent* variable t (“time”); superimposed dots indicate time differentiations; the symbol \cdot denotes the standard *scalar product* in the plane (so that, in particular, $r^2 = \vec{r} \cdot \vec{r}$ is the squared modulus of the 2-vector $\vec{r} \equiv \vec{r}_1 - \vec{r}_2$); and ρ_1 , ρ_2 , and ω are 3 *arbitrary real nonvanishing* parameters. It is shown below that this model is *Hamiltonian* and solvable by *algebraic* operations: in fact by just finding the 2 zeros of a second-degree time-dependent polynomial explicitly known in terms of the initial data of the problem, $\vec{r}_n(0)$ and $\dot{\vec{r}}_n(0)$. Moreover—if the 2 parameters ρ_1 , ρ_2 are both *rational* numbers—then this model is *isochronous*, its *generic* solutions being *all* periodic with a common period which is an *integer* multiple of the basic period $T = 2\pi/|\omega|$. Note that this model is invariant under a common multiplicative rescaling of the dependent variables $\vec{r}_n \equiv \vec{r}_n(t)$, hence it is in particular *invariant under rotations* of the Cartesian plane (as demonstrated by the covariance of its equations of motion (1.1)); while it is *not* translation-invariant. ■

Example 1.2. The equations of motion of the *second* of these 2 *solvable* models read as follows:

$$\begin{aligned} \ddot{\vec{X}} = & \omega \dot{\vec{X}} + \rho_1 \omega^2 \vec{X} + \frac{1}{X^2} \left[2 \left(\dot{\vec{X}} \cdot \dot{\vec{X}} \right) \dot{\vec{X}} - \left(\dot{\vec{X}} \cdot \dot{\vec{X}} \right) \dot{\vec{X}} \right], \\ \ddot{\vec{x}} = & -\frac{1}{2} (\rho_2 \omega)^2 \vec{x} - \frac{1}{x^2} \left\{ 2 \left(\dot{\vec{x}} \cdot \dot{\vec{x}} \right) \dot{\vec{x}} - \left(\dot{\vec{x}} \cdot \dot{\vec{x}} \right) \dot{\vec{x}} \right. \\ & \left. + \left[\rho_1 + (\rho_2)^2/2 \right] \omega^2 \left[2 \left(\vec{X} \cdot \vec{x} \right) \vec{X} - X^2 \vec{x} \right] \right\} \end{aligned} \quad (1.2a)$$

$$\begin{aligned}
 & + \omega \left[\left(\dot{\vec{X}} \cdot \vec{x} \right) \hat{X} + (\vec{X} \cdot \vec{x}) \dot{\hat{X}} - \left(\dot{\vec{X}} \cdot \dot{\vec{X}} \right) \hat{x} \right] \\
 & + \left[2 \left(\dot{\vec{X}} \cdot \vec{x} \right) \dot{\vec{X}} + \left(\dot{\vec{X}} \cdot \dot{\vec{X}} \right) \vec{x} \right] \}. \quad (1.2b)
 \end{aligned}$$

Here the 2 *dependent* variables $\vec{X} \equiv \vec{X}(t) \equiv (X_1(t), X_2(t))$ and $\vec{x} \equiv \vec{x}(t) \equiv (x_1(t), x_2(t))$ are 2-vectors describing the motion in the *real* Cartesian plane of 2 unit-mass point-particles as functions of the *independent* variable t (“time”); $\hat{X} \equiv \hat{X}(t) \equiv (-X_2(t), X_1(t))$, and $\hat{x} \equiv \hat{x}(t) \equiv (-x_2(t), x_1(t))$ are these vectors rotated in the Cartesian plane by $\pi/2$; $X \equiv X(t)$ and $x \equiv x(t)$ are the moduli of these 2-vectors; the symbol \cdot denotes the standard *scalar product* in the plane (so that, for instance, $x^2 = \vec{x} \cdot \vec{x} = \hat{x} \cdot \hat{x}$); superimposed dots indicate time-differentiations; and the 3 parameters ρ_1, ρ_2, ω are 3 *arbitrary real* parameters. Note that we used the *unambiguous* notation $\dot{\vec{X}} \cdot \dot{\vec{X}} \equiv (\dot{X}_1)^2 + (\dot{X}_2)^2$ (instead of the *ambiguous* notation $(\dot{X})^2$ that may rather be interpreted to mean $\left\{ d/dt [(X_1)^2 + (X_2)^2]^{1/2} \right\}^2 = [(\dot{X}_1 X_1 + \dot{X}_2 X_2)/X]^2$); and likewise for $\dot{\vec{x}} \cdot \dot{\vec{x}} \equiv (\dot{x}_1)^2 + (\dot{x}_2)^2$. Also note that these equations of motion are *scale-invariant*, i.e., *invariant* under the *rescaling* transformation $X(t) \Rightarrow c X(t)$, $x(t) \Rightarrow c x(t)$ with c an *arbitrary nonvanishing constant*; hence, they are *covariant*, i.e., *rotation-invariant* in the plane. And they are clearly *Newtonian*: accelerations equal forces, with the forces depending *nonlinearly* on the coordinates of the particles and on their velocities.

A *remarkable* feature of this system of 2 equations of motion (see (1.2)) is that—if the 2 parameters ρ_1 and ρ_2 are *rational* numbers,

$$\rho_n = \frac{q_n}{k_n} \quad (1.2c)$$

where q_n and k_n are two *arbitrary coprime integers* ($q_n \neq 0$, $k_n \geq 1$, $n = 1, 2$), and ω is an *arbitrary nonvanishing real* parameter to which the period $T = 2\pi/|\omega|$ is associated—then this model is *isochronous*: there is a *large open* set of initial data $\vec{x}(0), \dot{\vec{x}}(0), \vec{X}(0), \dot{\vec{X}}(0)$ yielding *nonsingular* evolutions that are *all periodic* with a period that is an *integer* multiple of the basic period $T = 2\pi/|\omega|$. ■

These findings are merely *rather trivial quite special* examples of those discussed later in the book (see Chapter 4, where however we generally consider N -body problems with *arbitrary* $N \geq 2$). The interested reader will find their proofs there, and thereby learn how to obtain the *explicit* solutions of the initial-value problems of these 2 Newtonian 2-body systems in the plane.