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J. Y. Buchanan

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COMPTES RENDUS
OF
OBSERVATION AND REASONING

BY

J. Y. BUCHANAN, M.A., F.R.S.

Commandeur de l'Ordre de Saint Charles de Monaco

Vice-Président du Comité de Perfectionnement de l'Institut Océanographique
(Fondation Albert I^{er} Prince de Monaco)

“Prove all things. Hold fast that which is good.”

I Thess. v. 21

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PREFACE

AS the title of this volume indicates, the book consists of “accounts rendered” of work done at different times, in different places and on different subjects.

In republishing papers, many of which are of almost ancient date, it was thought advisable to accompany them by explanatory notes and comments. As it was inconvenient to introduce these in the text, they have been embodied in the Contents, which form in fact a summary of the work.

The Contributions to Newspapers concerning matters of public interest at the time have been reprinted because they are of public interest still.

Of the scientific communications the most important are those concerning the Natural History of Steam and Ice, and these have been reprinted in their original form, although this has involved some reduplication of matter. This has been accepted on account of the importance of the experiments and of the consequences which follow from them and of the apparent unwillingness of the scientific public to make use of them.

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Preface

In the *Chemical and Physical Notes*, which formed part of the *Antarctic Manual* of 1901, will be found a résumé of my experiments in the domain of Inorganic Natural History which, from my own experience, I judged would be found useful by the Chemists and Physicists of the Expeditions for which the *Manual* was prepared.

It was conveyed to me through an old friend and former colleague that this contribution to the *Antarctic Manual* had done much to retard the Standardisation of Research. I took it as a compliment. To standardise research is to limit its freedom and to impede discovery. Originality and independence are the characteristics of genuine research, and it is stultified by the acceptance of standards and by the recognition of authority.

J. Y. BUCHANAN.

26 May 1916.

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This investigation was undertaken in order to determine experimentally whether the salt which occurs in the ice formed by freezing sea-water and other saline solutions is a part of the solid ice or belongs to the liquid brine which adheres to the ice. The fundamental principle on which the experiments were based is that, if a saline solution, such as sea-water, is partially frozen and the temperature of the mixture is observed, then, if the ice so formed is pure ice, it may be removed and be replaced by pure ice of independent origin, such as snow, having the same temperature, and if heat is then supplied, the snow will melt in the solution at the same temperature as that at which the ice which was produced in the solution was found to melt. This was found to be the case. As a corollary,

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it follows that the temperature at which pure ice melts depends on the medium in which it melts. The subject is dealt with under two heads ; namely : (a) the temperature at which sea-water and some other saline solutions freeze, and the chemical constitution of the solid and the liquid into which they are split by freezing ; and (b) the temperature at which pure ice melts in sea-water and in a number of saline solutions of different degrees of concentration	133
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This is the name which I invented for my own use in recording the occurrence of a common surface feature of tropical and equatorial regions, when on board the “Challenger.” I found in all such countries that the rocks were decomposed to a depth of many metres, the residual material often remaining in situ, with such a fresh appearance that it was difficult to imagine that it could be anything but unaltered rock. It was only necessary, however, to touch it with a stick or even with the fingers, for it to “crumble” into fragments of all sizes down to sand and clay. In almost every place within the tropics visited for the first time the rocks were logged as consisting of “the crumble formation.” Outside of the tropics this formation occurs only in a rudimentary form. The crumble formation owes its existence to a high atmospheric temperature and to humidity. It is the typical formation produced by the subaerial weathering of rock in situ. Very slight mechanical disturbance is sufficient to cause degradation under gravity, with the production, first of taluses ; then as the taluses undergo further gravitational degradation, they flatten out, and the final product of the crumble formation and of subaerial weathering is the pampa, prairie, steppe or desert, which are the native names for the same formation in homologous climatic regions of the earth.

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No. 7. BEOBACHTUNGEN ÜBER DIE EINWIRKUNG DER STRAH- LUNG AUF DAS GLETSCHEREIS. (Extract of paper read before the physical section of the Schweizerische Naturforschende Gesellschaft at its meeting at Basel on 6 September, 1910, and printed in its <i>Verhandlungen</i> , I, p. 330)	28c
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As almost all my work on ice, outside of the laboratory, was done in Switzerland, and as the Swiss Naturforschende Gesellschaft honoured me many years ago by making me one of its honorary members, I considered it my duty to communicate to it an account of the more important results of my work on the Swiss glaciers. I had also in my mind that in case of there being any doubt about the work, the Swiss were geographically in a better position to verify it than any other people. As the meeting was being held in the principal city of German Switzerland, I made my communication in the German language.	
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While engaged in discussing questions connected with the physics of the ocean, I found the want of definite knowledge of the amount of solar heat which really reaches the surface of the land or sea in a form which can be collected, measured and utilised. There was no lack of actinometrical observations, but I found it impossible from them to obtain the data that I sought. The aim of most observers has been to arrive, by more or less direct means, at what is known as the solar constant, that is, the quantity of heat which is received in unit time by unit surface, when exposed perpendicularly to the sun's rays outside of the limits of the earth's atmosphere. For my purpose the amount of radiation arriving at the outside of the earth's atmosphere was of no importance. What I wanted to know and to measure was the amount of solar radiation which strikes the earth at the sea-level and is there revealed as heat. It is the energy of this radiation which maintains the terrestrial economy. Having the opportunity of accompanying the expedition to Egypt for observing the total eclipse of the sun on May 17th, 1882, I determined to	

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<i>Locality.</i>	
The astronomers chose a spot on the banks of the Nile close to the town of Sohag, and in Latitude 26° 37' N. for the observation of the eclipse, and experience showed that it had been well chosen. The eclipse was total at 8.34 a.m. on the 17th May 1882, civil reckoning. The maximum duration of totality that was expected was 70 seconds, and, in fact, it lasted longer than 65 seconds. The expedition arrived on the 8th May and I began work with the calorimeter on the 11th and devoted the whole of my time to this work until the 19th May, when the camp was struck. From the outset the calorimeter worked most satisfactorily and the only alteration which had to be made was to replace the original metal dome, as steam space of the boiler, by a glass tube. From the 11th to the 15th May I was occupied in studying the instrument and learning how to use it, in the only way by which this is possible, namely, by setting it to do the work expected of it, noticing deficiencies of arrangement and mistakes made in handling, and rectifying them as they showed themselves. In the course of this educational work valuable preliminary results were obtained, and on the 16th, 17th and 18th trustworthy experiments were carried out with the apparatus in best working order and under very favourable natural conditions	338
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No. 12. THE TOTAL SOLAR ECLIPSE OF AUGUST 30, 1905. (From *Nature*, December 21, 1905, Vol. LXXIII, p. 173.) . . . 399

As this eclipse was to be visible in an easily accessible part of Europe, where the probability of finding fine weather was great and the calculated duration of totality was nearly four minutes, I determined to see it. A considerable display of protuberances was expected and I wished to form my own idea of their size by personally checking their persistence or non-persistence through the phase of totality as seen from a station situated as nearly as possible on the line of mid-totality. The display of protuberances at the moment of second contact was very brilliant; when the time of mid-totality arrived not a trace of them was visible to the naked eye. Therefore these very brilliant protuberances had an apparent height less than 45 seconds of arc . . . 400

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This remarkable eclipse had a peculiar interest of its own, and the observers of it had the advantage that the central line passed through Paris and many other important places in Northern Europe, but up to the last moment it was uncertain whether the eclipse would be total, annular or partial. The value accepted for the apparent diameter of the moon was therefore of paramount importance. I observed it in front of the school house at Euabonne, a northern suburb of Paris, and used an ordinary binocular when the naked eye was not enough, and a hand-screen made of three coloured glasses which reduced the density of the sun's light without altering its colour. This was a very efficient instrument and I am sorry that I do not know the source from which it was obtained. I bought it from a hawkler in the streets of Barcelona, on the eve of the eclipse of August 30, 1905, and of course, like all hawkler's goods, it was anonymous	405
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