

ACCRETION PROCESSES IN ASTROPHYSICS

It has been more than fifty years since the first significant paper on accretion flows was written. In recent years, X-ray satellites capable of identifying accretion disks and radiation jets – indications that accretion has taken place – have significantly advanced our understanding of these phenomena. This volume presents a comprehensive and upto-date introduction to the major theoretical and observational topics associated with accretion processes in astrophysics. Comprising lectures presented at the twenty-first Winter School of the Canary Islands Institute of Astrophysics, the text emphasizes the physical aspects of accretion, investigating how radiation jets are produced, how accretion power is divided between jets and radiated energy, the geometry of accretion flow, and the accretion processes of active galactic nuclei. Written by an international team of experienced scientists, chapters offer young researchers key analytical tools for supporting and carrying out the next generation of front-line research.

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978-1-107-03019-0 - Accretion Processes in Astrophysics: XXI Canary Islands Winter School of Astrophysics Edited by Ignacio González Martínez-País, Tariq Shahbaz and Jorge Casares Velázquez

Front matter

More information

Canary Islands Winter School of Astrophysics

Volume XXI

Accretion Processes in Astrophysics

Series Editor

Francisco Sánchez, Instituto de Astrofísica de Canarias

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ACCRETION PROCESSES IN ASTROPHYSICS

XXI Canary Islands Winter School of Astrophysics

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More information

CAMBRIDGE UNIVERSITY PRESS

University Printing House, Cambridge CB2 8BS, United Kingdom

Cambridge University Press is part of the University of Cambridge.

It furthers the University's mission by disseminating knowledge in the pursuit of education, learning and research at the highest international levels of excellence.

www.cambridge.org

Information on this title: www.cambridge.org/9781107030190

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First published 2014

A catalogue record for this publication is available from the British Library

Library of Congress Cataloging in Publication data

Canary Islands Winter School of Astrophysics (21st: 2009: Puerto de la Cruz, Canary Islands) Accretion processes in astrophysics / [edited by] Ignacio González Martínez-País, Instituto de Astrofísica de Canarias and Universidad de La Laguna, Tariq Shahbaz, Instituto de Astrofísica de Canarias, Jorge Casares Velázquez, Instituto de Astrofísica de Canarias.

pages cm

Lectures presented at the XXI Canary Islands Winter School of Astrophysics, held in Puerto de la Cruz, Tenerife, Spain, Nov. 2–13, 2009.

Includes bibliographical references.

ISBN 978-1-107-03019-0 (hardback)

1. Accretion (Astrophysics) – Congresses. I. González Martínez-País, Ignacio, 1959 – editor of compilation. II. Shahbaz, Tariq, 1970 – editor of compilation. III. Casares Velázquez, Jorge, 1964 – editor of compilation. IV. Title.

QB466.A25C36 2009

523.01-dc23 2012051629

ISBN 978-1-107-03019-0 Hardback

Cambridge University Press has no responsibility for the persistence or accuracy of URLs for external or third-party internet websites referred to in this publication, and does not guarantee that any content on such websites is, or will remain, accurate or appropriate.



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Frontmatter More information

Contents

Lis	st of contributors	page ix
Lis	st of participants	xi
Pr	eface	xiii
Ac	knowledgments	XV
Αŀ	obreviations	xvii
1	Accretion disks Henk Spruit	1
2	The evolution of binary systems Philipp Podsiadlowski	45
3	Accretion onto white dwarfs Brian Warner	89
4	Multiwavelength observations of accretion in low-mass X-ray binary systems $Robert\ I.\ Hynes$	117
5	X-ray binary populations in galaxies Giuseppina Fabbiano	151
6	Observational characteristics of accretion onto black holes I $\it Christine\ Done$	184
7	Observational characteristics of accretion onto black holes II: environment and feedback ${\it Rob~Fender}$	227
8	Computing black-hole accretion John F. Hawley	253
A	Piazzi Smyth, the Cape of Good Hope, Tenerife, and the siting of large telescopes Brian Warner	291





978-1-107-03019-0 - Accretion Processes in Astrophysics: XXI Canary Islands Winter School of Astrophysics Edited by Ignacio González Martínez-País, Tariq Shahbaz and Jorge Casares Velázquez

Frontmatter More information

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Frontmatter

More information

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xi



978-1-107-03019-0 - Accretion Processes in Astrophysics: XXI Canary Islands Winter School of Astrophysics Edited by Ignacio González Martínez-País, Tariq Shahbaz and Jorge Casares Velázquez

Frontmatter

More information

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Preface

It was more than 50 years ago when the first significant paper on accretion flows was written. Since then, the subject has grown incredibly, and today many X-ray satellites are engaged in research into observational signatures and tests of theoretical models for accretion processes in astrophysics. Recognizing the continued importance of this field, the Instituto de Astrofísica de Canarias organized the XXIst in its Winter School series around the topic "Accretion Processes in Astrophysics."

The primary aim of the school was to provide a wide-ranging and up-to-date overview of the theoretical, experimental, and analytical tools necessary for carrying out front-line research in the study of accretion processes. The school was particularly designed to offer young researchers guidelines to support their research in these areas.

The 40 lectures presented a fairly comprehensive and up-to-date introduction to the major observational and theoretical topics associated with accretion. With emphasis on the physical processes involved, this includes applications to close binary systems such as cataclysmic variables and X-ray binaries and their evolution, as well as the theory of relativistic accretion flows and the accretion processes in active galactic nuclei. The lectures were given by eight experienced scientists who are actively working on a variety of leading research projects and who have played key roles in the advances made in the field in recent years.

The Editors





Acknowledgments

The editors would like to express their warmest gratitude to all the lecturers for their time in preparing their classes, attending the school, and writing the chapters for this book. We know that this has been a major effort on their part, but we hope that it has been rewarding for them. In particular we would like to thank Prof. John F. Hawley for his entertaining public lecture on "Black Holes" and Prof. Brian Warner for his ad hoc lecture dedicated to Charles Piazzi Smyth, who built the first (temporary) major observatory on Tenerife.

The key to the success of the school has been without any doubt our secretary Lourdes González. Without her help and diligence, the school would not have worked as smoothly as it did. We also thank Nieves Villoslada, who started the preparation and organization of the school; Jesús Burgos of the OTRI, who provided invaluable help with the preparation of applications needed to receive funding; persons at the IAC's Centro de Cálculo for their IT assistance; and Ismael Martínez Delgado, for the technical editing of this book for Cambridge University Press.

We are extremely grateful to the local artist Macu Anelo, who designed the school's poster that depicts an accretion disk around a black hole, superimposed on a background of Guanches sketches, which we hope may entice young scholars to enter this field. The school's poster was prepared by Ramón Castro. We thank Annia Domènech and the Gabinete de Dirección of the IAC for taking the time to conduct interviews and submit press releases for all the lecturers.

We greatly acknowledge the financial assistance from the Spanish Ministerio de Educación y Ciencia and from the Cabildo de Tenerife, who kindly provided the excellent facilities of the Congress Palace of Puerto de la Cruz where the event took place. Last, but not least, we would like to acknowledge all the participants of the school: lecturers, students, and supporting personnel.

The Editors





978-1-107-03019-0 - Accretion Processes in Astrophysics: XXI Canary Islands Winter School of Astrophysics Edited by Ignacio González Martínez-País, Tariq Shahbaz and Jorge Casares Velázquez

Frontmatter

More information

Abbreviations

AAVSO American Association of Variable Star Observers

ACF Auto Correlation Function

ACIS AXAF CCD Imaging Spectrometer ADAF Advection Dominated Accretion Flows

ADC Accretion Disk Corona AGB Asymptotic Giant Branch AGN Active Galactic Nucleus

ASCA Advanced Satellite for Cosmology and Astrophysics

AU Astronomical Unit BAL Broad Absorption Line

BATSE Burst And Transient Source Experiment

BB Black Body BH Black Hole

BHB Black Hole Binaries

BHXRT Black Hole X-Ray Transient

BLR Broad Line Region

BPS Binary Population Synthesis CCD Charge Coupled Device CCF Cross Correlation Function

CE Common Envelope

CGRO Compton Gamma-Ray Observatory

CT Constrained Transport CV Cataclysmic Variable

CXB Cosmological X-ray Background

DD Double Degenerate
DIM Disk Instability Model

DN Dwarf Nova DNe Dwarf Novae

DNO Dwarf Nova Oscillation
DNS Double Neutron Star
EMF Electromotive Force
FIR Far Infra Red

FRED Fast Rise Exponential Decay

GC Globular Cluster GRB Gamma Ray Burst

GRMHD General Relativistic Magneto Hydrodynamics

HID Hardness Intensity Diagram HMXB High Mass X-ray Binary HST Hubble Space Telescope

IAC Instituto de Astrofísica de CanariasIMBH Intermediate-Mass Black HoleIMXB Intermediate-Mass X-ray Binary

IP Intermediate Polar

IR Infrared

ISAF Ion Supported Accretion Flow ISCO Innermost Stable Circular Orbit

ISM Interstellar Medium

IUE International Ultraviolet Explorer

KG Kilo Gauss

LARPS Low Accretion Rate Polars

xvii



978-1-107-03019-0 - Accretion Processes in Astrophysics: XXI Canary Islands Winter School of Astrophysics Edited by Ignacio González Martínez-País, Tariq Shahbaz and Jorge Casares Velázquez

Front matter

More information

xviii	Abbreviations		
LGRB LIGO LINER LL	Long-Duration Gamma Ray Burst Laser Interferometer Gravitational-wave Observatory Low Ionization Nuclear Emission-line Region Landau and Lifshitz (1959)		
$_{ m LMC}$	Large Magellanic Cloud		
LMXB	Low-Mass X-ray Binary		
LOFAR	Low-Frequency Array		
lpDNO	longer-period Dwarf Nova Oscillation		
MG	Mega Gauss		
MHD	Magneto Hydrodynamics		
MRI	Magneto Rotational Instability		
MWA	Murchison Widefield Array		
NASA	National Aeronautics and Space Administration		
NLR	Narrow Line Region		
NS	Neutron Star		
NSE	Nuclear Statistical Equilibrium		
NTT	New Technology Telescope		
PPM	Piecewise Parabolic Method		
PS	Population Synthesis		
PSF	Point Spread Function		
QPO RLOF	Quasi Periodic Oscillation Roche Lobe Overflow		
$egin{array}{c} { m ROSAT} \\ { m RXTE} \end{array}$	Röentgen Satellite		
SALT	Rossi X-ray Timing Explorer Southern African Large Telescope		
SALI	Single Degenerate		
sdB	subdwarf Binary		
SED	Spectral Energy Distribution		
SFR	Star Formation Rate		
SLE	Shapiro-Lightman-Eardley solutions		
SPY	SN Ia Progenitor SurveY		
SS	Shakura-Sunyaev		
STIS	Space Telescope Imaging Spectrograph		
SXT	Soft X-ray Transient		
SyS	Symbiotic Star		
TZO	Thorne Żytkow Object		
UCXB	Ultracompact X-ray Binary		
ULX	Ultra Luminous X-ray source		
UV	Ultraviolet		
VLBI	Very Long Baseline Interferometry		
WD	White Dwarf		
WRLOF	Wind Roche Lobe Overflow		
XLF	X-ray Luminosity Function		
XMM	European Space Agency's X-ray Multi-mirror Mission		
VDD	V D D'		

XRB

X-Ray Binary