

# ASTRONOMY

## THE EVOLVING UNIVERSE

### NINTH EDITION

**Michael Zeilik**  
University of New Mexico



**CAMBRIDGE**  
UNIVERSITY PRESS

[www.cambridge.org](http://www.cambridge.org)

## About the Author

Dr. Michael Zeilik works as Professor of Astronomy at the University of New Mexico. In his teaching, he specializes in innovative, introductory courses for the novice, non-science major student. His classes include cooperative learning teams to explore key astronomical concepts with hands-on activities. He has been supported by grants from the National Science Foundation, NASA, the Exxon Educational Foundation, and the Slipper Fund of the National Academy of Sciences for innovations in astronomy education, delivery of astronomy to the general public, and astronomy workshops for in-service teachers.

Dr. Zeilik's current research activities focus on two areas: astronomy education and astronomy in the historic and prehistoric Pueblo world in the U.S. Southwest. He has published more than 100 professional articles and four books and has given more than 200 talks to professional and lay audiences.

Dr. Zeilik earned his A.B. in Physics with honors at Princeton University and his M.A. and Ph.D. in Astronomy at Harvard University. He has been a Woodrow Wilson Fellow, a National Science Foundation Fellow, and a Smithsonian Astrophysical Observatory Predoctoral Fellow. At the University of New Mexico, he has been named a Presidential Lecturer, the highest award for all-around performance by a faculty member. In 1998, he was appointed a Research Fellow with the National Institute of Science Education.

In 2000, he became Chair of the Astronomy Education Committee of the American Association of Physics Teachers.

Dr. Zeilik is listed in *American Men and Women of Science*, *The Writers Directory*, *Contemporary Authors*, *Who's Who in the West*, and *Who's Who of Emerging Leaders in America*. He is a member of the Authors Guild and the Text and Academic Authors Association. The 8th edition of *Astronomy: The Evolving Universe* won a 1997 Texty Award from the Text and Academic Authors Association.



Contents

Preface	xi	Chapter 2 The Birth of Cosmological Models	22
Concept Clusters	xv	2.1 Scientific Models	24
How to Study Astronomy	xvii	<i>Building Models</i>	24
		<i>Evaluation of Models</i>	24
		2.2 Greek Models of the Cosmos	25
		<i>Harmony and Geometry</i>	25
		<i>A Physical Geocentric Model</i>	26
		<i>A Contrary View: A Sun-Centered Model</i>	27
		<i>Stellar Parallax in a Finite Cosmos</i>	27
		<i>Expanding the Geocentric Model</i>	29
		2.3 Claudius Ptolemy:	
		A Total Geocentric Model	31
		A Geocentric Model Refined	31
		Nonuniform Motion	32
		Ptolemy's Complete Model	34
		The Size of the Cosmos	34
		Model Evaluation	35
		2.4 Describing Basic Observations	
		with a Geocentric Model	35
		Enrichment Focus 2.1	
		Surveying the Earth	28
		Chapter 3 The New Cosmic Order	40
		3.1 Copernicus the Conservative	42
		<i>The Heliocentric Concept</i>	43
		<i>The Plan of De Revolutionibus</i>	43
		3.2 The Heliocentric Model of Copernicus	44
		<i>Details of the Model</i>	44
		<i>Retrograde Motion Explained Naturally</i>	45
PART 1 Changing Conceptions of the Cosmos	1		
Chapter 1 From Chaos to Cosmos	2		
1.1 The Visible Sky	4		
<i>Constellations</i>	4		
<i>Angular Measurement</i>	4		
<i>Motions of the Stars</i>	4		
1.2 The Motions of the Sun	7		
<i>Motions Relative to the Horizon</i>	8		
<i>Motions Relative to the Stars</i>	9		
<i>View from the Southern Hemisphere</i>	12		
1.3 The Motions of the Moon	12		
<i>Motions Relative to the Stars</i>	12		
<i>Phases</i>	13		
1.4 The Motions of the Planets	14		
<i>Retrograde Motion</i>	14		
<i>Elongations, Conjunctions,</i>			
<i>and Oppositions</i>	15		
<i>Angular Speed and Relative Distances</i>	16		
1.5 Eclipses of the Sun and Moon	16		
Enrichment Focus 1.1			
Precession of the Equinoxes	11		
Enrichment Focus 1.2			
Angular Size and Speed	18		

<i>Planetary Distances</i>	46	<i>Orbits and Escape Speed</i>	84
<i>Relative Distances of the Planets</i>	46	<i>Newton's Cosmology</i>	85
<i>Problems with the Heliocentric Model</i>	48	<b>Enrichment Focus 4.1</b>	
<i>The Impact of the Heliocentric Model</i>	49	Speed, Velocity, and Acceleration	69
3.3 Tycho Brahe: First Master of Astronomical Measurement	50	<b>Enrichment Focus 4.2</b>	
<i>The New Star of 1572</i>	50	Newton, the Apple, and the Moon	80
<i>Tycho's Hybrid Model</i>	50	<b>Enrichment Focus 4.3</b>	
3.4 Johannes Kepler and the Cosmic Harmonies	52	The Mass of the Sun	81
<i>The Harmonies of the Spheres</i>	51	<b>Enrichment Focus 4.4</b>	
3.5 Kepler's New Astronomy	53	Escape Speed	86
<i>The Battle with Mars</i>	53	<b>Chapter 5 The Birth of Astrophysics</b>	90
<i>Properties of Ellipses</i>	54	5.1 Sunlight and Spectroscopy	92
<i>Kepler's Laws of Planetary Motion</i>	55	<i>Atoms and Matter</i>	92
<i>The New Astronomy</i>	57	<i>A Model of the Atom</i>	92
<b>Enrichment Focus 3.1</b>		<i>Simple Spectroscopy</i>	93
Sidereal and Synodic Periods in a Heliocentric Model	47	5.2 Analyzing Sunlight	94
<b>Enrichment Focus 3.2</b>		<i>Kirchhoff's Rules</i>	96
Geometry of Ellipses	54	<i>The Conservation of Energy</i>	97
<b>Chapter 4 The Clockwork Universe</b>	62	<i>Kinetic Energy</i>	98
4.1 Galileo: Advocate of the Heliocentric Model	64	<i>Potential Energy</i>	99
<i>The Magical Telescope</i>	64	5.3 Spectra and Atoms	99
<i>The Starry Messenger</i>	66	<i>Light and Electromagnetic Radiation</i>	99
<i>Galileo's Discoveries and the Copernican Model</i>	66	<i>Waves</i>	99
<i>The Crime of Galileo</i>	66	<i>The Electromagnetic Spectrum</i>	100
4.2 Galileo and a New Physics of Motion	68	<i>Atoms, Light, and Radiation</i>	101
<i>Acceleration, Velocity, and Speed</i>	68	<i>Solving the Puzzle of Atomic Spectra</i>	102
<i>Natural Motion Revisited</i>	69	<i>Energy Levels</i>	103
<i>Forced Motion: Gravity</i>	70	<i>Other Atoms</i>	105
<i>Planets and Pendulums</i>	70	5.4 Spectra from Atoms	105
<i>Galileo's Cosmology</i>	71	<i>Absorption-Line Spectra</i>	106
4.3 Newton: A Physical Model of the Cosmos	72	<b>Enrichment Focus 5.1</b>	
<i>The Prodigious Young Newton</i>	72	Kinetic Energy	98
<i>The Magnificent Principia</i>	73	<b>Enrichment Focus 5.2</b>	
<i>Forces and Motions</i>	73	Energy and Light	102
<i>Newton's Laws of Motion</i>	74	<b>Chapter 6 Telescopes and Our Insight into the Cosmos</b>	110
4.4 Newton and Gravitation	76	6.1 Observations and Models	112
<i>Centripetal Acceleration</i>	76	6.2 Visible Astronomy: Optical Telescopes	113
<i>Newton's Law of Gravitation</i>	78	<i>The Basis of Optics</i>	113
4.5 Cosmic Consequences of Universal Laws	81	<i>Optics and Images</i>	114
<i>The Earth's Rotation</i>	81	<i>Telescopes</i>	115
<i>Precession of the Earth's Axis</i>	81	<i>Functions of a Telescope</i>	116
<i>The Earth's Revolution and the Sun's Mass</i>	82	<i>Next Generation of Telescopes</i>	118
<i>Gravity and Orbits</i>	82	<i>New Optical Techniques</i>	119
		6.3 Invisible Astronomy	120
		<i>Ground-Based Radio</i>	121
		<i>Resolving Power and Radio Interferometers</i>	122



vi • Contents

<i>Ground-Based Infrared</i>	123	<i>Origin and Development of the Oceans</i>	160
<i>Space Astronomy</i>	124	<i>Evolution of the Atmosphere</i>	161
6.4 Image Collection and Processing	125	<i>Evolution of the Earth's Surface Temperature</i>	161
<i>Understanding Intensity Maps</i>	125	8.7 The Earth's Evolution: The Big Picture	161
<i>Photography</i>	126	8.8 Models of Planets	163
<i>Charge-Coupled Devices (CCDs)</i>	126	<b>Enrichment Focus 8.1</b>	
<b>Enrichment Focus 6.1</b>		Radioactivity and the Dating of Rocks	154
Properties of Telescopes	117		
<b>Chapter 7 Einstein's Vision</b>	130		
7.1 Natural Motion Reexamined	133	<b>Chapter 9 Moon and Mercury, Mars and Venus: Terrestrial Planets</b>	166
<i>Newton's Assumptions</i>	133	9.1 General Orbital and Physical Characteristics	168
<i>Motion and Geometry</i>	133	<i>Moon</i>	168
7.2 The Rise of Relativity	134	<i>Mercury</i>	170
<i>Mass and Energy, Space and Time</i>	134	<i>Mars</i>	172
<i>The General Theory of Relativity</i>	135	<i>Venus</i>	172
<i>The Principle of Equivalence</i>	135	9.2 Surface Environments	174
<i>Weightlessness and Natural Motion</i>	136	<i>Moon</i>	174
7.3 The Geometry of Spacetime	137	<i>Mercury</i>	179
<i>Euclidean Geometry</i>	137	<i>Mars</i>	180
<i>Non-Euclidean Geometry</i>	138	<i>Venus</i>	183
<i>Local Geometry and Gravity</i>	138	<i>The Active Surface of Venus</i>	184
<i>The Curvature of Spacetime</i>	139	9.3 Magnetic Fields	186
<i>Spacetime Curvature in the Solar System</i>	139	<i>Moon</i>	186
<i>Experimental Tests of General Relativity</i>	141	<i>Mercury</i>	187
7.4 Geometry and the Universe	142	<i>Mars</i>	188
<i>Cosmic Geometry</i>	142	<i>Venus</i>	188
7.5 Relativity and the Cosmos	143	9.4 History and Evolution	189
<i>Escape Speed and the Critical Density</i>	143	<i>Moon</i>	189
<i>The Future of the Universe</i>	143	<i>Mercury</i>	189
		<i>Mars</i>	191
		<i>Venus</i>	192
<b>PART 2 The Planets: Past and Present</b>	147	<b>Enrichment Focus 9.1</b>	
<b>Chapter 8 The Earth: An Evolving Planet</b>	148	Tides and Tidal Friction	169
8.1 The Mass and Density of the Solid Earth	150	<b>Enrichment Focus 9.2</b>	
8.2 The Earth's Interior and Age	150	Distances in the Solar System	171
8.3 The Earth's Magnetic Field	153	<b>Enrichment Focus 9.3</b>	
<i>Origin</i>	155	Impact Cratering	177
<i>Magnetosphere</i>	155		
8.4 The Blanket of the Atmosphere	155	<b>Chapter 10 The Jovian Planets: Primitive Worlds</b>	196
<i>The Greenhouse Effect</i>	156	10.1 Jupiter: Lord of the Heavens	198
<i>Atmospheric Circulation</i>	157	<i>Physical Characteristics</i>	200
8.5 The Evolution of the Crust	157	<i>Atmospheric Features and Composition</i>	200
<i>Planetary Evolution and Energy</i>	158	<i>A Model of the Interior</i>	202
<i>Volcanism and Plate Tectonics</i>	158	<i>Magnetic Field</i>	203
8.6 Evolution of the Atmosphere and Oceans	159		

10.2 The Many Moons and Rings of Jupiter	203	<i>Models of Origin</i>	241
<i>Io</i>	204	11.5 Basics of Nebular Models	241
<i>Europa</i>	205	<i>Angular Momentum</i>	241
<i>Ganymede</i>	205	<i>Heating of the Nebula</i>	244
<i>Callisto</i>	206	11.6 The Formation of the Planets	244
<i>The Rings of Jupiter</i>	207	<i>Making Planets</i>	244
10.3 Saturn: Jewel of the Solar System	207	<i>Chemistry and Origin</i>	245
<i>Atmosphere and Interior</i>	207	<i>Jupiter and Saturn: A Different Story?</i>	247
<i>Similarities to Jupiter</i>	208	<i>Our Moon: A Different Story!</i>	248
10.4 The Moons and Rings of Saturn	209	<i>Evaluation of the Nebular Model</i>	248
<i>Titan</i>	211	Enrichment Focus 11.1	
<i>Other Moons</i>	211	Momentum and Angular Momentum	242
<i>Ring System</i>	213		
10.5 Uranus: The First New World	214		
<i>Atmospheric and Physical Features</i>	214	<b>PART 3 The Universe of Stars</b>	<b>251</b>
<i>Moons and Rings</i>	215		
<i>Magnetic Field</i>	216	<b>Chapter 12 Our Sun: Local Star</b>	<b>252</b>
10.6 Neptune: Guardian of the Deep	217	12.1 A Solar Physical Checkup	254
<i>Physical Properties</i>	217	<i>How Far?</i>	254
<i>Moons and Rings</i>	217	<i>How Big?</i>	254
<i>Atmospheric Features</i>	220	<i>How Massive?</i>	254
<i>Magnetic Field</i>	220	<i>How Dense?</i>	254
10.7 Pluto and Charon: Guardians of the Dark	221	12.2 Ordinary Gases	255
<i>Orbital and Physical Properties</i>	221	<i>Temperature</i>	255
<i>Charon: Pluto's Companion Planet</i>	222	<i>Pressure</i>	256
Enrichment Focus 10.1		12.3 The Sun's Continuous Spectrum	256
The Doppler Shift	201	<i>Luminosity</i>	256
		<i>Surface Temperature and Blackbody Radiators</i>	257
		<i>Opacity</i>	260
<b>Chapter 11 The Origin and Evolution of the Solar System</b>	<b>226</b>	12.4 The Solar Absorption-Line Spectrum	260
11.1 Debris between the Planets: Asteroids	228	<i>The Origin of Absorption Lines</i>	260
<i>Asteroids: Minor Planets</i>	228	<i>The Chemical Composition of the Photosphere</i>	261
<i>Composition</i>	229	12.5 Energy Flow in the Sun	262
11.2 Comets: Snowballs in Space	229	<i>Conduction, Convection, Radiation</i>	262
<i>Composition</i>	232	<i>Photosphere</i>	262
<i>Orbits</i>	233	<i>Chromosphere</i>	263
<i>The Comet Cloud</i>	233	<i>Corona</i>	263
<i>Interactions with the Solar Wind</i>	233	<i>Solar Wind</i>	265
<i>Halley's Comet</i>	234	12.6 The Solar Interior	265
<i>Comet Shoemaker-Levy 9 and the Great Comet Crash on Jupiter</i>	237	<i>Energy Sources</i>	265
11.3 Meteors and Meteorites	237	<i>Nuclear Transformations</i>	265
<i>Types of Meteorites</i>	238	<i>Fusion Reactions</i>	266
<i>Origin of Meteorites</i>	238	<i>Solar Neutrino Problem</i>	268
11.4 Pieces and Puzzles of the Solar System	240	<i>Solar Vibrations and Interior</i>	268
<i>Chemistry</i>	240	12.7 The Active Sun	269
<i>Dynamics</i>	240	<i>Sunspots</i>	269
		<i>Sunspot Cycle</i>	269
		<i>Physical Nature of Sunspots</i>	271
		<i>Flares</i>	271
		<i>Coronal Loops and Holes</i>	272

viii • Contents

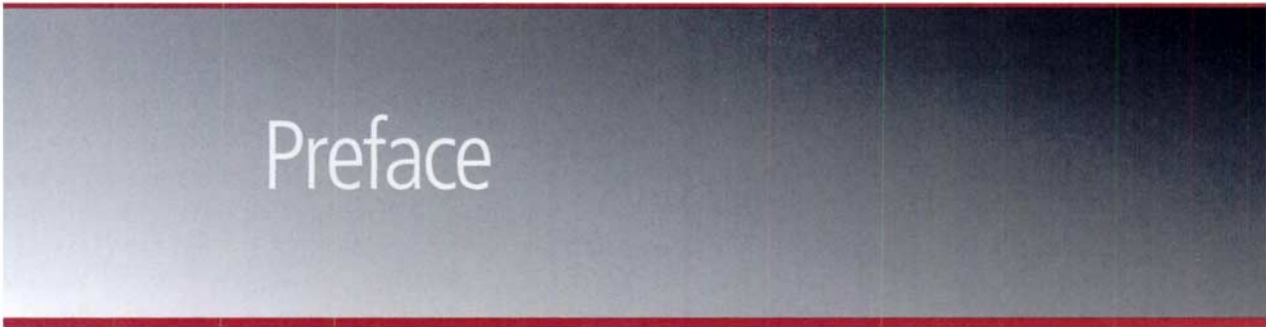
Enrichment Focus 12.1			
The Sun's Luminosity	257		
Enrichment Focus 12.2			
Emission from Blackbodies	259		
<b>Chapter 13 The Stars as Suns</b>	<b>278</b>		
13.1 Some Messages of Starlight	280		
Brightness and Flux	281		
Flux and Luminosity	281		
The Inverse-Square Law for Light	283		
13.2 Stellar Distances: Parallaxes	284		
13.3 Stellar Colors, Temperatures, and Sizes	287		
Color and Temperature	287		
Temperature and Radius	287		
Direct Measurement of Diameters	289		
13.4 Spectral Classification of Stars	289		
Temperature and the Balmer Lines	289		
Spectral Classification	290		
13.5 The Hertzsprung–Russell Diagram	292		
Temperature versus Luminosity	293		
Spectroscopic Distances	295		
13.6 Weighing and Sizing Stars:			
Binary Systems	296		
Binary Stars	296		
Spectroscopic Binary Systems	298		
The Mass-Luminosity Relation			
for Main-Sequence Stars	298		
Stellar Densities	299		
Stellar Lifetimes	300		
Enrichment Focus 13.1			
Flux and Magnitude	282		
Enrichment Focus 13.2			
Heliocentric (Trigonometric)			
Stellar Parallax	286		
<b>Chapter 14 Starbirth and Interstellar Matter</b>	<b>304</b>		
14.1 The Interstellar Medium: Gas	306		
Bright Nebulas	306		
Interstellar Atoms	306		
21-cm Emission from Atomic Hydrogen	307		
Clouds and Intercloud Gas	310		
Interstellar Molecules	310		
Molecular Clouds	311		
14.2 The Interstellar Medium: Dust	312		
Cosmic Dust	312		
Dust and Infrared Observations	315		
The Nature of Interstellar Dust	315		
Dust and the Formation			
of Molecules	316		
Formation of Cosmic Dust	317		
14.3 Starbirth: Theoretical Ideas	317		
Collapse Models	317		
Protostar Formation	317		
Collapse with Rotation	318		
14.4 Starbirth: Observational Clues	319		
Signposts for the Birth			
of Massive Stars	320		
The Birth of Massive Stars	320		
The Birth of Solar-Mass Stars	320		
Molecular Outflows and Starbirth	322		
Planetary Systems?	323		
Brown Dwarfs: Failed Stars	325		
14.5 Neighboring Planetary Systems:			
Found!	326		
Center-of-Mass Motions	326		
Doppler Shift Detections	327		
Enrichment Focus 14.1			
Emission Nebulas: Forbidden Lines	310		
<b>Chapter 15 Star Lives</b>	<b>332</b>		
15.1 Stellar Evolution and the			
Hertzsprung–Russell Diagram	334		
Classifying Objects	334		
Time and the H–R Diagram	335		
15.2 Stellar Anatomy	336		
Pressure and Energy Balance	336		
Energy Transport	336		
15.3 Star Models	337		
Goals of Models	338		
General Results	338		
15.4 Energy Generation and the			
Chemical Compositions of Stars	338		
Hydrogen Burning	338		
Helium and Carbon Burning	339		
15.5 Theoretical Evolution			
of a 1-Solar-Mass Star	340		
Evolution to the Main Sequence	340		
Evolution on the Main Sequence	341		
Evolution off the Main Sequence	342		
Evolution to the End	343		
The Fate of the Earth	344		
Lower-Mass Stars	344		
Chemical Composition and Evolution	344		
15.6 Theoretical Evolution			
of Massive Stars	344		
Evolution of a 5-Solar-Mass Star	345		
More Massive Stars	345		
15.7 Observational Evidence			
for Stellar Evolution	346		
Stars in Groups	346		
Globular Clusters	347		
Stellar Populations	349		



<i>Comparison with the H–R Diagram of Clusters</i> 350	<b>PART 4 Galaxies and Cosmic Evolution</b> 387
<i>Variable Stars</i> 351	<b>Chapter 17 The Evolution of the Galaxy</b> 388
<i>Central Stars of Planetary Nebulas</i> 352	17.1 The Galaxy’s Overall Structure 390
15.8 The Synthesis of Elements in Stars 353	17.2 Galactic Rotation: Matter in Motion 392
<i>Nucleosynthesis in Red Giant Stars</i> 353	<i>The Sun’s Speed around the Galaxy</i> 392
<b>Enrichment Focus 15.1</b>	<i>The Sun’s Distance from the Center</i> 392
<i>Degenerate Gases</i> 342	<i>Rotation Curve and the Galaxy’s Mass</i> 393
<b>Chapter 16 Stardeath</b> 358	17.3 Galactic Structure from Optical Observations 395
16.1 White Dwarf Stars: Common Corpses 360	<i>Spiral Tracers and Spiral Structure</i> 395
<i>Physics of Dense Gases</i> 360	<i>Optical Maps of Spiral Structure</i> 396
<i>White Dwarfs in Theory</i> 360	17.4 Exploring Galactic Structure by Radio Astronomy 397
<i>Observations of White Dwarfs</i> 361	<i>Radial Velocities and Rotation</i> 397
16.2 Neutron Stars: Compact Corpses 361	<i>Radio Maps of Spiral Structure</i> 397
<i>Degenerate Neutron Gases</i> 361	17.5 The Evolution of Spiral Structure 398
<i>Physical Properties</i> 362	<i>The Windup Problem</i> 398
16.3 Novas: Mild Stellar Explosions 362	<i>The Density-Wave Model</i> 399
<i>Ordinary Novas</i> 363	<i>Status of the Density-Wave Model</i> 400
<i>A Nova Model</i> 363	17.6 The Heart of the Galaxy 401
16.4 Supernovas: Cataclysmic Explosions 364	<i>Radio Observations</i> 401
<i>Classifying Supernovas</i> 365	<i>Infrared Observations</i> 402
<i>The Origin of Supernovas</i> 366	<i>X-Rays</i> 402
<i>Supernova 1987A</i> 367	<i>The Inner 30 Light Years</i> 403
<i>Supernova Remnants</i> 368	<i>Does a Black Hole Lurk in the Core?</i> 403
<i>The Crab Nebula:</i>	17.7 The Halo of the Galaxy 404
<i>A Supernova Remnant</i> 371	<i>Globular Clusters</i> 404
16.5 The Manufacture of Heavy Elements 373	<i>Other Material in the Halo</i> 405
<i>Nucleosynthesis in Stars</i> 373	17.8 A History of Our Galaxy 405
<i>Nucleosynthesis in a Supernova</i> 373	<i>Populations and Positions</i> 405
<i>Other Sites of Nucleosynthesis</i> 374	<i>The Birth of the Galaxy</i> 406
16.6 Pulsars: Neutron Stars in Rotation 374	<b>Enrichment Focus 17.1</b>
<i>Observed Characteristics</i> 375	<i>The Mass of the Galaxy</i> 394
<i>Clock Mechanism</i> 375	<b>Chapter 18 The Universe of Galaxies</b> 410
<i>Pulsars and Supernovas</i> 375	18.1 The Extragalactic Debate 412
<i>A Lighthouse Model for Pulsars</i> 376	<i>Basic Arguments</i> 412
<i>Binary Radio Pulsars</i> 377	<i>The Debate’s Resolution</i> 413
<i>Very Fast Pulsars</i> 377	18.2 Normal Galaxies: A Galaxian Zoo 414
<i>Pulsars with Planets</i> 378	<i>Ellipticals</i> 415
16.7 Black Holes: The Ultimate Corpses 378	<i>Disks (Spirals)</i> 415
<i>The Schwarzschild Radius</i> 379	<i>Irregulars</i> 416
<i>The Singularity</i> 380	<i>Luminosity Classes</i> 416
<i>Perilous Journey into a Black Hole</i> 380	18.3 Surveying the Universe of Galaxies 418
16.8 Observing Black Holes 381	<i>Judging Distances</i> 418
<i>Binary X-Ray Sources</i> 381	<i>Distance Indicators</i> 418
<i>Observing Black Holes</i> 382	18.4 Hubble’s Law and Distances 419
16.9 High-Energy Bursters Everywhere! 383	<i>Redshifts and Distances</i> 419
<i>X-Ray Bursters</i> 383	<i>“Age” of the Cosmos</i> 422
<i>Gamma-Ray Bursters: Mystery Solved?</i> 383	<i>An Accelerating Expansion?</i> 423
<b>Enrichment Focus 16.1</b>	
<i>Thermal and Nonthermal (Synchrotron) Emission</i> 370	

x • Contents

18.5	General Characteristics of Galaxies	423	Enrichment Focus 19.1	
	Size	423	Faster Than Light?	456
	Mass	424		
	Luminosities	425	Chapter 20 Cosmic History	462
	Mass–Luminosity Ratios	425	20.1 Cosmological Assumptions	
	Colors	426	and Observations	464
	Spin	427	Assumptions	464
18.6	Clusters of Galaxies	427	Observations	466
	The Local Group	427	A Brief Review of Cosmology	466
	Other Clusters of Galaxies	429	20.2 The Basic Big Bang Model	466
	Clusters and the Luminosity of Galaxies	429	20.3 The Cosmic Background Radiation	467
	Interacting Galaxies	430	Discovery!	467
18.7	Superclusters and Voids	433	Confirmation!	468
	The Cosmic Tapestry	434	Properties	470
18.8	Intergalactic Medium		20.4 The Primeval Fireball	470
	and Dark Matter	434	The Hot Start	470
	Enrichment Focus 18.1		Creation of Matter from Photons	471
	Hubble’s Constant and the Age		Temperature Greater than $10^{12}$ K	472
	of the Universe	422	Temperature about $10^{12}$ K	472
			Temperature about $10^9$ K	472
			Temperature about 3000 K	473
			Evidence for the Big Bang	473
Chapter 19	Cosmic Violence	438	20.5 The End of Time?	474
19.1	Violent Activity in Galaxies	440	20.6 From Big Bang to Galaxies	475
	Evidence of Violence in Our Galaxy	440	20.7 Elementary Particles and the Cosmos	476
	Synchrotron Emission Revisited	441	The Forces of Nature	476
	Active Galaxies	441	GUTs and the Cosmos	477
19.2	Radio Galaxies	442	GUTs and Galaxy Formation	478
	Two High-Profile Active Galaxies	443	20.8 The Inflationary Universe	479
	Structures of Radio Emission	445	The Flatness Problem	479
19.3	Seyfert Galaxies and BL Lacertae		The Horizon Problem	480
	Objects	446		
	Seyfert Galaxies	446	Appendix A Units	485
	BL Lacertae Objects	448	Appendix B Planetary Data	489
19.4	Quasars: Unraveling the Mystery	449	Appendix C Physical Constants	
	Quasar Redshifts	450	and Astronomical Data	494
	General Observed Properties	450	Appendix D Nearby Stars in the <i>Hipparcos</i>	
	The Light from Quasars	451	Catalogue	495
	Line Spectra	451	Appendix E The Most Luminous Stars	
	Variability in Luminosity	452	in the <i>Hipparcos</i> Catalogue	496
19.5	Grand Illusions: Gravitational Lenses	452	Appendix F Periodic Table of the Elements	497
19.6	Troubles with Quasars	453	Expanded Glossary	499
	Energy Sources in Quasars	455	Index	533
	A Generic Quasar and AGN Model	457		
	A Model for Quasars and AGNs	457		
	The Host Galaxies for Quasars	459		



A *ninth* edition! I really cannot believe that an idea that I had some twenty-five years ago has lived so long and gone so far. Many instructors and students have told me that they’ve enjoyed using the book, and that feedback has certainly helped me along. I also desire to improve this book so that it innovates and evolves as a better learning tool. Finally, astronomy changes rapidly, especially with the advent of new space and ground-based telescopes. Our outer vision grows, and our inner one intensifies. So what’s new for this edition?

One, the obvious **updating of the material**. Constant change drives the excitement of astronomy. That is a main reason why astronomy appeals to me, and, I hope, to you. There’s always something new and unexpected for us to discover in the universe. However, I have done the updating with great care to focus on discoveries that I think have long-term value and connect with the major concepts.

Two, a **streamlining of the material** so as to make the descriptions, concepts, and explanations as clear and concrete as possible. I have taken great care to minimize the use of technical terms and the passive voice. I also gave special attention to the Learning Outcomes and Key Concepts to keep them concise and precise.

Three, a **refined art program**. The figures have been reviewed and revised with the goal of

better understanding by the novice student (while keeping the science accurate). I have tried to incorporate a human perspective whenever possible. I have also aimed for clarity and simplicity so that careful “viewing” of the figures will illuminate key concepts.

Four, a **research-based pedagogy**. I have targeted concepts in this edition more than ever, based on research in astronomy education carried out by myself and others. But concepts in isolation have little value – they must be *connected*. The right connections of key concepts result in the “big picture” of the cosmos. My goal for my course and this book is to provide a connected understanding of astronomy. To do so, I have introduced a new organizing feature, Celestial Navigators™.

Each chapter contains one **Celestial Navigator™**. These maps provide visual guides of major concepts in the chapter and explicitly show their connections. I have tested them in my classes, and they result in large, robust gains in students’ connected understanding of astronomy.

To promote an even higher level of conceptual unification, each part begins with a **Part Concept**, followed by an **Inquiry Focus** – a series of “How do we know . . . ?” questions to link concepts across chapters within a Part. Again, the goal is the “big picture,” this time within a thematically-unified Part.

THEMATIC AND TOPICAL STRUCTURE

Previous editions have developed the notion of “how do we know?” in astronomy. I have made that question more explicit and the central theme for the entire book. I believe that answering the “how do we know?” promotes the understanding of any subject, but astronomy more so than many others do. Why? Because once we leave the earth and solar system, astronomy becomes quite abstract, essentially light, physics, and models.

Yet people tend to view science as a disconnected collection of facts, and our culture reinforces this view. Take for instance, the quiz show *Who Wants to Be a Millionaire*. The questions use a multiple-choice format, which is common in schools in the United States. Here is one question for the big prize, \$1,000,000:

How long does it take the light from the sun to reach the earth?

- A. 42 seconds
- B. 3 minutes
- C. 8 minutes
- D. 1 hour

Do you know the answer? (It is C.) Now, that’s OK, but do you know the unasked question? *How do we know that the sun is 8 light-minutes from the earth?* That’s a much deeper question and requires a level of understanding far more fundamental than the first (which can just be memorized).

**“How do we know?” is the governing theme of this edition.** In terms of topics, this book has two main aspects: to describe in narrative form the range of the astronomical universe and its cosmic connections; and to introduce students to how astronomers think about the cosmos so they can gain some understanding about its operation. I hope that students will become so enticed by the contents that they will become intrigued by the concepts linking and illuminating astronomical phenomena.

This edition is designed for a one-semester introduction to astronomy. Note that it is much shorter than previous editions, with fewer and shorter chapters. Yet it retains the previous structure of four coherent parts, each focusing on a key subtheme of cosmic evolution. Like the cosmos, each part connects to the others, so you can really approach the four parts in any order.

Part 1: Changing Conceptions of the Cosmos

This part accents the evolution of cosmological ideas, from the geometric views of the Greeks to the mind-boggling visions of modern astronomy. It leads off with the simplest observations you can make without a telescope from the earth and ends with the farthest reaches of the visible universe. Part 1 acquaints the reader with the idea of *scientific models*, the conceptual core of modern scientific thought. Scientific models are born from our imagination and experience; they mark the essential creative act of the scientific enterprise. As such models evolve, they shape our changing conceptions of the cosmos. The development of scientific models resounds throughout the book; it is *the* fundamental tool to understand the universe.

Essential to the formation of models are new astronomical observations. In the age of microelectronics and space astronomy, we have greatly expanded our vision of the cosmos. Many space telescopes, especially the Hubble Space Telescope, marked great changes in the outer vision of twentieth-century astronomy. In this century, we know that new technology telescopes on the earth will surpass space telescopes in some ways. We can hope that lunar-based telescopes will augment the observational legacy of space telescopes. New generations of computers endow our minds with another way to see the physical processes in the cosmos by the simulation of astronomical systems. Together, computers and telescopes work as the tools to impart innovative ideas to astronomers.

Part 2: The Planets: Past and Present

Flyby spacecraft and gangly landers have provided new insights to our understanding of the planets. This part focuses on the physical properties of the planets to infer their origin and evolution. It first takes a comparative look at our current knowledge of the planets, especially our earth and moon, as well as Mercury, Venus, and Mars. These planets show different degrees of evolution, with the earth being the most evolved. The other planets – Jupiter, Saturn, Uranus, Neptune, and Pluto – have, in contrast, changed little since their birth.

Space missions have disclosed that the moons of the outer planets are really worlds unto themselves, places of rock and ice scarred by violence in the

past. These new worlds provide important clues, along with comets and meteorites, of the early history of the solar system. So planetary evolution traces back to origin – the birth of the solar system from an interstellar cloud of gas and dust. By astronomical standards, that birth was quick, violent, and chaotic – shaping the primitive forms of the planets. This picture implies that many other single stars have planetary systems and that these other worlds resemble the local planets in broad ways.

Part 3: The Universe of Stars

Our sun and its planets swing around in a vast island of stars called the Milky Way Galaxy – our home galaxy. As the nearest star, our sun serves as the close-up model for other stars, especially to understand the physical processes within them. The Galaxy contains some hundreds of billions of stars at various stages of their lives. These stars have been born, like the sun, from clouds of interstellar gas and dust. Modern technology supplies us with views deep into the regions of starbirth, showing us the early lives of stars.

Because stars live long by human standards, we cannot directly observe their evolution. We can build models of stars with computers, and these models provide us with the dimension of time to map out the lives of stars. Ordinary stars, powered by fusion reactions, grow old and blow off material before their demise. Many have violent deaths, leaving bizarre corpses such as neutron stars. Massive stars undergo violent deaths, which are signaled by enormous explosions that build heavy elements and propel them into space to seed the next generation of stars. The span of the lives of stars guides us to an understanding of what will happen to our sun in its old age.

Part 4: Galaxies and Cosmic Evolution

The universe contains galaxies, in which the most visible material of the cosmos resides. This part first examines ordinary galaxies, like our Milky Way, and then hyperactive ones. We now realize that galaxies with unusual activity are interacting by gravity with their neighbors, which triggers the celestial fireworks. The show includes long, thin jets of material, confined by magnetic fields, rocketing out of the cores of these galaxies. All these galaxies

dwell in clusters with other galaxies, and we have just come to the realization that clusters are laid out in long chains with vast voids in between – the cosmic tapestry. Amid this remarkable layout of visible matter lurks matter that we cannot yet see – the so-called dark matter that shapes the visible universe, both its structure and destiny.

The cosmic design in the large-scale architecture was imprinted in the awesome explosion in which the universe began. That explosion – the Big Bang – linked the smallest pieces of matter to the universe at large. The Big Bang also left relics that we observe today, evidence of the violence of creation. From the Big Bang some 15 billion years ago, cosmic evolution shaped us at where we are and when we are in the universe.

QUICK START

This book is for students, who are probably novices to astronomy. I want them to learn effectively from it. I have designed a four-part structure so that you can investigate each part somewhat independently of the others. Many cross-references, especially to the basic physical and astronomical ideas, should help you to link the parts together. I have made a concerted effort to introduce ideas as concretely as possible. Within parts, I deal with the most familiar first: Chapter 1 (Part 1) with the visible sky; Chapter 8 (Part 2) with the earth; Chapter 12 (Part 3) with the sun; and Chapter 17 (Part 4) with the Milky Way Galaxy. Within chapters, I have tried to present concrete examples before abstract notions.

The **Enrichment Focus** sections furnish another linkage throughout the text. I have set these optional sections off from the main text to enrich ideas by basic mathematics (algebra, trigonometry, and geometry). You will need to decide which of them will be specifically assigned. You will note problems and activities at the end of each chapter. Some of them draw on the material in the Focus sections.

To aid novice science students, I worked to simplify the language as much as I can by using ordinary English rather than technical jargon. My rule when using technical terms is: *Define every term you use, and use every term you define.* I try to avoid a one-time use of a technical term just for the sake of completeness.

If your students are really baffled by a good



xiv • Preface

strategy to learn astronomy, please tell them to read carefully the special section by Mark Hollabaugh (who teaches at Normandale Community College in the United States) on “How to Study Astronomy.” It provides a general strategy for studying plus specific guides to each of the learning features of this book.

**NOTE.** If you have picked up this book because you are curious about astronomy, you may be interested in taking a college-level course by mail for academic credit. Write to: Independent Study, Continuing Education, The University of New Mexico, 1634 University NE, Albuquerque, NM 87131, USA. The course is called Astronomy 101C; I am the instructor.

INSTRUCTORS’ RESOURCES

We have centralized all instructor resources on the Web site for this book. The URL is

<http://www.TheEvolvingUniverse.com>.

ACKNOWLEDGMENTS

Locally, Sabrina Moore initiated the effort to acquire the new visual images in the book. Louise Shaler provided in-depth editorial work on the manuscript.

She also assisted me in the final round of exploring for astronomical images and acquiring the permissions for them. Artist Boris Starosta and I collaborated on the art and the creation of the Celestial Navigator™ maps.

Other folks include my correspondence students (who have just the book and their brains) and the students in my regular Astronomy 101 course at UNM. I have also received letters and e-mail from students at other colleges and universities, as well as from instructors. Thank you all!

Any errors in the text are my responsibility. It is a little known fact that minor corrections and changes *can* be made in future printings of *this* edition. Please keep this fact in mind and send me any errors you may find. When ordering, please request the *latest printing* of the book so your students will have the most correct version.

Your feedback can improve this book! Please send any comments to me at the Department of Physics and Astronomy, The University of New Mexico, 800 Yale Blvd NE, Albuquerque, NM, 87131-1156, USA. My Internet mail address:

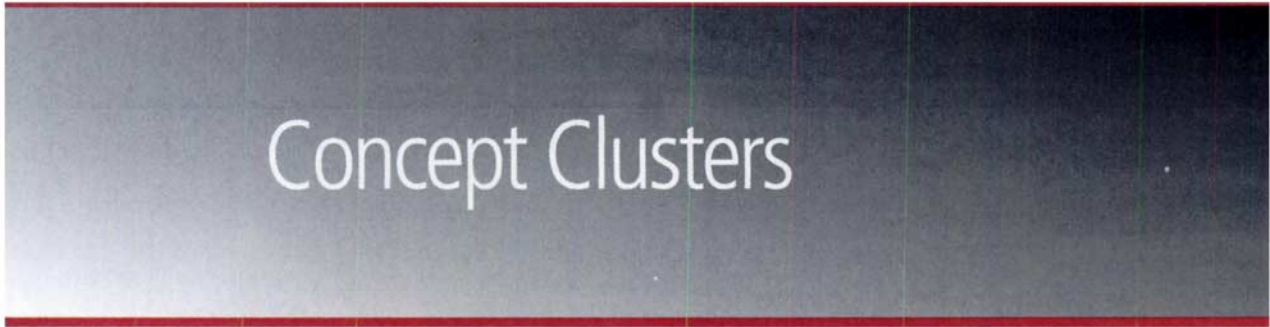
[zeilik@la.unm.edu](mailto:zeilik@la.unm.edu).

**Michael Zeilik**  
Santa Fe, New Mexico, U.S.A., October 2001

Abbreviations

Abbreviations are often used for the names of major observatories and agencies (particularly in the figure captions). These are:

AAO – Anglo-Australian Observatory	NCSA – National Center for Supercomputer Applications
AUI – Associated Universities, Inc.	NOAO – National Optical Astronomy Observatories
AURA – Association of Universities for Research in Astronomy	NRAO – National Radio Astronomy Observatory
CTIO – Cerro Tololo Interamerican Observatory	NSO – National Solar Observatory
ESA – European Space Agency	SAO – Smithsonian Astrophysical Observatory
ESO – European Southern Observatory	SOHO – Solar and Heliospheric Observatory
HST – Hubble Space Telescope	STScI – The Space Telescope Science Institute
KPNO – Kitt Peak National Observatory	TRACE – Transition Region and Coronal Explorer
NASA – National Aeronautics and Space Administration	VLA – Very Large Array radio telescope
NCAR – National Center of Atmospheric Research	VLT – Very Large Telescope, ESO



The material in this book falls into four clusters of related concepts; these clusters themselves are interrelated. The Celestial Navigators(tm) include a subset of the most important concepts, also organized by cluster.

**1. Cosmic Distances**

- Angles/angular diameters/positions
- Angular speeds/relative motions
- Astronomical Unit/Kepler’s laws
- Heliocentric parallax/triangulation
- Inverse-square law for light/flux
- Luminosity classes from H-R diagram/  
spectroscopic distances
- Doppler shifts of interstellar clouds/rotation  
curve of the Galaxy
- Period-luminosity relation for cepheids
- Structure of the Milky Way Galaxy/distances
- Distance indicators to galaxies (cepheids,  
supernovas)
- Clusters and superclusters of galaxies
- Hubble law/Hubble constant/red shifts/  
age of universe

**2. Heavenly Motions**

- Angles/angular speeds/relative distances
- Motions of sun, moon relative to horizon,  
stars/eclipses

- Motions of the planets relative to sun, stars/  
retrogrades/oppositions/elongations
- Geocentric/heliocentric
- Heliocentric parallax
- Kepler’s laws/orbits/periodic motion/  
dark matter
- Newton’s laws of motion and gravitation/  
orbits/mass/weight/freefall/escape speed
- Binary stars/masses of stars/center of mass
- Tidal forces
- Gravitational accretion/contraction
- General relativity/spacetime/curved geometry
- Conservation of energy/types of energy
- Conservation of angular momentum

**3. Celestial Light and Spectra**

- Electromagnetic radiation/spectrum
- Emission/absorption
- Kirchhoff’s rules (emission, absorption, contin-  
uous spectra)
- Atomic energy levels/photons/excitation/  
radiative energy

Telescopes: detectors; resolving, light-gathering, magnifying power; interferometers  
Planck curve/black body/colors/temperature  
Synchrotron emission/magnetic fields  
Fusion reactions/nucleosynthesis  
Energy transport (radiation, convection, conduction)  
Chemical composition/spectra  
Doppler shift/radial velocity/blue and red shifts  
Stellar spectral and luminosity classes  
Hertzsprung-Russell (H-R) diagram/star clusters  
Mass-luminosity (M-L) relation/stellar lifetimes  
Spectra of the interstellar medium/intergalactic medium  
Spectra of galaxies/ red shifts/expansion of cosmos  
Cosmic background radiation/Big Bang model

4. Scientific Models

Assumptions, aesthetics/Geometry, physics  
Observations (errors)/Predictions, explanations  
Quantum theory/photons/energy levels  
Age/radioactive dating/half life  
Properties of matter (solids, liquids, gases, plasmas)/density  
Solar system (geocentric/heliocentric)  
Planets/planetary evolution/tectonics/volcanism/thermal energy  
Magnetic fields/dynamo model  
Sun/stars/stellar evolution/H-R diagram  
Novas/supernovas/nucleosynthesis/mass loss  
Stellar corpses /white dwarfs/neutron stars/black holes  
Starbirth/planetary systems/protostars and protoplanets/brown dwarfs  
Milky Way Galaxy/normal galaxies  
Active galaxies/supermassive black holes  
Formation of galaxies/protogalaxies  
Big Bang model/critical density/inflationary models/GUTs

# How to Study Astronomy

**Mark Hollabaugh**  
*Normandale Community College, Minnesota, U.S.A.*

I'd like to give you some suggestions to assist your learning.

Welcome to astronomy! When I was growing up in Michigan, I would go outside on crisp winter nights to watch the northern lights dance across the sky. In August, I counted meteors during the Perseid meteor shower. I could imagine the terror ancient peoples must have felt as the moon turned blood red during a lunar eclipse. I looked through a real telescope for the first time when I was in seventh grade. A science teacher who spent summers in our town set up his telescope each Friday night in the city park. After a brief peek through the scope, I would run back to the end of the line for another look. I was hooked on astronomy. Later, after studying physics, I did research on comets and asteroids, and on cosmic rays.

Throughout these experiences, I have marveled with fascination at the grandeur and intricacies of the universe. Now, I try to pass that fascination on to my students. That is also what Dr. Zeilik does in *Astronomy: The Evolving Universe*. My association with Mike began with my interest in Native American astronomy. We also share an interest in helping college students like you to enjoy learning astronomy effectively. Your professor has chosen a superb textbook for you to use in your astronomy course.

## UNDERSTANDING YOUR LEARNING STYLE

Each of us learns in different ways. Some of us learn best by reading. Others find listening to a lecture to be most helpful. Talking with others can help us clarify our own understanding. Drawing diagrams or pictures can help us see the relationships among concepts.

When learning something new, it helps to know our own strengths, weaknesses, and learning style. As you begin your study of astronomy, make a list of the kinds of activities that help you learn best. You will find a variety of learning resources presented in your astronomy course: lectures, textbook readings, laboratory exercises, classroom demonstrations, computer simulations, learning team activities, Web sites, videotapes, or slides. These are all a part of your "tool kit" for understanding astronomy. Concentrate extra effort on those kinds of learning activities where you are the weakest. You may, for example, need to learn how to use your textbook as a resource to clarify ideas you don't understand from a lecture.

Memorization may help you correctly answer questions on an examination, but you will not be able to apply the concepts later. Avoid memorizing

## xviii • How to Study Astronomy

---

and seek rather to understand concepts and relate them to one another. The connected understanding of astronomy will reward you with the grand “big picture” of the cosmos!

### Learning to Learn

Ask yourself some questions about your study habits. Do you set aside a specific block of time each day to study a given subject? Do you work in a quiet place, free of distractions, such as the library? Do you study at a time of the day when you are at your best for learning new material? When you prepare for an examination, have you kept up with your reading of the text so all you need to do is review? Do you attend all class lectures and other activities? Do you concentrate on understanding concepts as opposed to memorizing facts or data? If you have answered yes to these questions, you have developed some excellent study habits! If not, set goals for yourself to change your learning habits during the current semester. Find out if your institution has an office that helps students to improve their study skills.

### Learning from Lectures

Should I read the book first or go to the lecture first? How you personally answer the question depends on your own learning style. It is usually helpful if you have some basic introduction to a topic before you go to your lecture. Look at the syllabus and make a note of the topics for the day. Then scan the *Learning Outcomes* for the chapter. Read the *Key Concepts* at the end of the appropriate chapter. Then skim the chapter in the textbook. If, for example, the lecture will be on Venus and Mars, page through Chapter 9. Examine the images and make note of any questions you might have about them. Pay special attention to the diagrams; be sure that you understand them. After class you should, of course, read the chapter more carefully.

### Learning by Taking Notes

Taking effective notes is a two-step process. First, you will make some notes in class as your professor lectures. Try to keep your notes in outline form. Don't worry about making a formal outline. Rather,

just indent subtopics under major topics. If your professor has provided a course outline for you, *use it*. Concentrate on getting down the “big ideas” and explanations of concepts. Your instructor wants you to *understand* astronomy! You can fill in the details later as you read the text.

The second step in taking good notes is editing them. Check for concepts or learning objectives that you don't understand and look them up in the textbook. Jot down any definitions you may have missed. In short, edit your notes, making sure you understand the principal ideas. This process will keep you well prepared for taking examinations.

### Learning with Other Students

Other students are *the* most important resource to your learning. You will understand more about astronomy if you explain, elaborate, or defend ideas to others. Forming small, informal study groups out of class is an important strategy. You can use the resources of the textbook to quiz one another. Come to such study groups prepared: Edit your notes, review the key concepts, and make a list of questions for your group. Learning together means you learn more. Don't wait to work together until the night before an examination. Make time in your schedule for regular meetings of your study group. If your institution has a computer network available to all students, you can contact other students by electronic mail or an electronic bulletin board.

Learning astronomy with a small group of students is an effective and fun way to increase your understanding of this fascinating subject. Your professor may use formal cooperative learning groups or may simply ask you to work together on a task or assignment. In either case following a few simple guidelines will enable your group to work well together. First, and perhaps most important, come to your group *prepared*. Be ready not only by mastering concepts in astronomy but also with questions about things that you do not yet understand.

Actively talk with one another. Make sure everyone participates. Ask questions. No idea or question is too trivial! Don't rush to conclusions, explore all the possibilities, and dig deeper! You will learn more in a learning group if you and the other group members explain or elaborate upon ideas. This process often happens in response to the simple



question, “Why?” So don’t be hesitant to challenge or to be skeptical about a group member’s idea. Finally, ask yourselves, “What did we do well as a group and what can we do better next time?” Working together will give you a better understanding of the evolving universe.

### Learning from the Textbook

Dr. Zeilik has incorporated many useful learning tools into your textbook to help you discover the excitement of astronomy. You should be aware of what your instructor expects of you. He or she probably will omit some topics, even whole sections of the textbook, perhaps even entire chapters! You may or may not be assigned specific Learning Outcomes from each chapter. It may help you to view your textbook as your primary reference. It supplements the lecture and other activities in the class. The textbook is your teacher away from the classroom. Your professor will expect you to read carefully and understand the assigned sections in *Astronomy: The Evolving Universe*.

Take notes as you read. Simply highlighting text is not particularly effective in placing concepts into our long-term memory. You will learn more, and recall it later, if the process of jotting down notes stimulates your brain as you read. Make notations in the margins of your textbook. Later, transcribe these notes into outline form to your notebook. Try to focus on the big ideas and reasons the universe works the way it does.

Take a few moments to become familiar with the design of the book by looking at the various sections of a chapter.

**Learning Outcomes.** Your instructor may indicate to you the learning outcomes for which you are responsible. These outcomes will help you know what is important. Before beginning to study the chapter, read the objectives. Then, as you read the chapter, think of the outcomes. *After* you have studied the chapter, write out a brief (one or two sentences) response to *each* outcome. If the outcome says “explain” or “describe,” do just that. Although this may seem time-consuming, it is an excellent way to summarize and apply what you have learned. You will find review questions at the end of each chapter keyed directly to the outcomes.

**Central Concept.** Each chapter has a theme. Keeping this theme in mind as you read will help you to understand the unifying idea of each chapter. Try to relate concepts back to the central question.

**Celestial Navigators™.** To help you see the “big picture,” each chapter contains one Celestial Navigator™ map. These serve as visual guides to the main concepts of a chapter and their connections with other concepts. You can copy these maps and use them as an organizing guides to the entire chapter. Write in your annotations on them as you read and study the material.

**Key Concepts.** When you read a mystery novel, you probably wouldn’t read the ending first. However, in reading *Astronomy: The Evolving Universe*, begin a chapter by studying the *Key Concepts* at the end of the chapter. Along with the Concept Clusters and Celestial Navigators™, this will give you an overview of the principal ideas in the chapter. The *Key Concepts* also will be very useful when reviewing for an examination.

**Enrichment Focus.** For some astronomy students, encountering a mathematical equation in a text can be both distracting and discomforting. You will notice that mathematical topics are discussed in short, compact essays. If your instructor covers these topics, you will want to review carefully these focus boxes. In most cases you should actually work through the calculations with your calculator. By using some of the data tables in the text, you can make up other examples to work.

**Study Exercises.** At the end of each chapter are a series of review exercises. They require you to explain concepts you have encountered in the chapter. Use these in your study group. These review exercises are keyed back to the Learning Outcomes. Test your understanding of specific outcomes with specific questions.

The most effective way to use these Study Exercises is to work on them with other students. You, and your friends, will learn more in the process. If you do not answer a question to your satisfaction, check the objective and discuss the material again. Then attempt the question a second time. It is not particularly beneficial for you simply to look up the answers to the multiple-choice questions. Try to give explanations of *why* your answer is correct, or *why* other answers to multiple-choice questions are incorrect.

xx • How to Study Astronomy

---

*Problems and Activities.* These typically require some knowledge of simple geometry or algebra. Or they may involve graphing or some observations. Your instructor will inform you about them.

*Key Terms.* Astronomy, and all of science, is rich with specific, technical terms. Each chapter includes boldface key terms. Write a brief definition of each term as you encounter it and then check it with the *Expanded Glossary*. Simply being able to define a term may not mean you understand it. Always give explanations. When you read, you may even encounter nontechnical words you don't fully understand. If you can't determine the meaning from the context, look the word up in a dictionary. A small, paperback dictionary in your book bag is as essential as a pen or calculator. If you forget the

meaning of an astronomical term, look in the *Expanded Glossary* first. The *Expanded Glossary* contains additional terms that may not be used in the text itself but which you may encounter in other astronomical readings.

*Appendices.* At the end of the textbook, these contain many tables of useful data and information. If it has been some time since you have taken a science class, or you are unfamiliar with the units of measurement in astronomy, you should carefully review Appendix A.

We hope you will enjoy your study of astronomy. Approach your class with a sense of awe and curiosity. Try to understand how the universe works. And, as another famous "science officer" has said, "Live long and prosper!"